# rtl-sdr and GNU Radio w/Realtek RTL2832U, E4000 and R820T



Originally meant for television reception and streaming the discovery and exploitation of the separate raw mode used in FM reception was perhaps first noticed by Eric Fry in March of 2010 and then expanded upon by Antti Palosaari in Feb 2012 who found that these devices can output unsigned 8bit I/O samples at high rates.

- librtlsdr contains the major part of the driver
   rtl\_sdr, rtl\_tcp, rtl\_test, etc command line capture tools and utilities
- gr-osmosdr gnuradio compatible module
   and a bunch of other stuff.

keenerd is the author of many other rtl\_\* tools: rtl\_fm, rtl\_power (heatmap.py), rtl\_adsb and code changes accepted into the mainline.

patchvonbraun is the author and maintainer of the build-gnuradio script that made it easy for me, and multitudes of others, to get started with rtlsdr under GNU Radio

### RF, DSP, and USB details

The dongles with an E4000 tuner can range between 54-2147 MHz (in my experience) with a gap over 1100-1250 MHz in general. The R820T and R820T2 go from 24-1760 MHz (but with reduced performance above 1500 MHz). The R820T dongles use a 3.57 MHz or 4.57 MHz intermediate frequency (IF) while the E4000s use a IQ pair Zero-IF. For both kinds the tuner error is ~30+20 PPM, relatively stable once warmed up, and stable from day to day for a given dongle. All of the generic dongle antenna inputs are 75 Ohm impedance (some SDR brande versions have 50 Ohm imput). The RTL2832 ADC differential input impedance is ~3,300 Ohm. The dynamic range for most dongles is around 45 dB. The sensitivity is somewhere around -110 dBm typically. The highest safe sample rate is 2.56 MS/s but in some situations up to 3.2 MS/s works without USB dropping samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL3820 U drops them internally). Because the devices use complex samples (RTL

The rtlsdr RTL2832U chips use a phased locked loop based synthesizer to produce the local oscillator required by the quadrature mixer. The quadrature mixer produces a complex-baseband output where the signal spans from -bandwidth/2 to +bandwidth/2 and bandwidth of the mixer output stages. (Datasheets, general refs: But what is the Fourier Transform? A visual introduction by 3BluelBrown, Quadrature Signals: Complex, But Not Complicated by Richard Lyons) This is complex-sampled (I and Q) by the ADC. The Sigma-Delta ADC samples at some high rate but low precision. From this a 28.8 Maps stream at 8 bits is produced. That can be resampled inside the RTL2832U to present whatever sample rate is desired to the host PC. This resampled output can be up to 3.2 MS/s but 2.5 MS/s is the max recommended to avoid losing samples. The minimum resampled output is 6.5 MS/s. Check this reddiff thread for caveadu output is interleaved; so one byte, I then one byte (P with no header or metadata (Intersatings). The samples themselves are unsigned and you subtract 127 from them to get their actual {-127,+127} value. You'll almost certainly notice a stable spike around DC. It's from either the 1/f noise of the electronics or if it's a Zero-IF tuner (E4000) the LO beating with itself in the mixer.

My favorite way to explore the spectrum is using rl. power to do very wideband multi-day surveys. For general use SDR# is probably the best application for windows with secondary mone-based linux and Mac support. I normally use Garx but it requires GNU Radio dependencies. Luckily there are Linux and OS X native binaries packages with all dependencies (ie, gurardio) these days. For doing diagnostic and low signal level work Linux and SX native Binaries packages with all dependencies (ie, gurardio) these way to use gefosphor, a GPU accelerated display, multimode has a very full and configurable CPU (It works great with GPU accelerated displays like ge-fosphor). For command line and low power devices ty keener's rl. fin.

Assuming you're on linux, but applicable in general, do not use the OS DVB drivers. Those are for the DVB-T mode and not the debug mode that outputs raw samples. Linux 3.x kernel should check with "\$ Ismod | grep dvb\_usb\_rtl28xxu" and if found at least "\$ sudo modprobe-r dvb\_usb\_rtl28xxu" to unload it.

While the sampling bandwidth is only 2.56 MHz the frequency can be re-tuned up to ~40 times a second. With frequency hopping you can survey very large bandwidths. See tholin's annotated 24 hour rtl\_power spectrogram. Below is a zoomable 37200\*31008 pixel 5 day long spectrogram I made using rtl\_power's FFT mode and heatmap.py. It starts very far zoomed out. It might load a bit slow too. (view full window)

This page is mostly just notes to myself on how to use rtlsdr's core applications, 3rd party stuff using librtlsdr and wrappers for it, and lots on using the gr-osmosdr source in GNU Radio and GNU Radio Companion. This isn't a "blog", don't read it sequentially, just search for terms of interest or use the topics menu. For realtime support on the same topics try Freenode IRC's ##rtlsdr and reddit's virtlsdr.



These days for most people doing most things you want to get an dongle with an R820T2 tuner. They'll come with MCX coaxial connectors. On sites like eBay shipping from China the average price is about ~\$10 shipped. These work fine for most things. At a bit higher price of ~\$20 some come with improvements like SMA or F connectors, metal cases and heatsinks on the tuner for stability above 1500 MHz, temperature controlled crystal oscillators, extra breakouts on the PCB, and the like.

I bought two E4000 based rtlsdr usb dongles for \$20 each in early 2012. Then many months later I bought two more R820T tuner based dongles for ~\$11 each. There's photos of the E4ks up at the top of the page and of an R820T based dongle in the "mini" format off to the left (most minis do not have ceproms for device ID). It and the Newsky E4k dongle up top are MCX. Back in 2012 some of the cheaper dongles occasionally miss protection diodes but that is no longer an issue. The antenna connector on the E4k ezcap up top is IEC-169-2, Belling-Lee. I usually replace it with an F-connector or use a PAL Male to F-Connector Female. F to MCX for the other style dongles. The default design has the tuner taking 75 Ohm so that's what they all are except SMA.

# Tuners

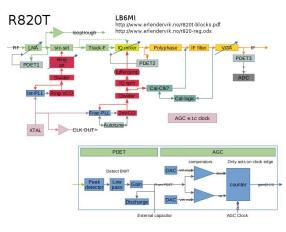
RTL-SDR Tuner Type Frequency Range 54 - 2200 MHz (1100 MHz-1250 MHz gap)
24 - 1766 MHz (>1500 MHz is bad w/o tuner cooling)
24 - 1766 MHz (>1500 MHz is bad w/o tuner cooling)
13 - 1864 MHz (mutabhilty's driver)
22 - 1100 MHz (FC0013BAC, FC00136 has seperate L band is
22 - 948 MHz
146 - 308 MHz and 438 - 924 MHz Elonics E4000 (E4K) Rafael Micro R820T Rafael Micro R820T2 Rafael Micro R820T2 Fitipower FC0012

Only three tuners are very desirable at this time. The Elonies E4000 and the Raphael Micro R820T/R820T2. In general they are of equal performance but the sticks with R820T2 chips are easy to find, cheaper (~\$10 USD), and they have a smaller DC spike due to the use of a non-zero intermediate frequency but must have cooling for the tuner to PLL lock above ~1500 MHz or so. The E4K is better for high end (~1.7.6Hz) while the R820T can tune down to 13 MHz without any hardware most (mutability's driver). The tuners themselves are set up and retuned with 12C commands. E4000 tuners used to re-tune twice as fast as R820T tuners, but this was fixed in keepended separational branch where R820T cally tune a limit to E44Ks. These changes were later adopted by the main risker.

For a detailed comparison of the E4000 vs R820T see the article, Some Measurements on E4000 and R820 Tuners: Image Rejection, Internal Signals, Sensitivity, Overload, 1 dB Compression, Intermodulation by HB9AJG in 2013.

Here's a dump of all the leaked datasheets and research by rtlsdr community members on the E4000 and R820T tuners I have as of 2020.

Стр. 1 из 21 20.11.2020, 23:11



- R820T datasheet-Non R-20111130 unlocked.pdf
- R8201 datasneet-Non R-20111130 uniocked.pdf http://www.erlendervik.no/r820-reg.ods LB6MI (local) http://www.erlendervik.no/r820t-blocks.pdf (local) http://prkele.prk.tky.fi/~peltolt/820T.xls (local) R820T2 Register Description.pdf

- R820T registers attempt.xls
- gat3way / r820tweak SDRSharp.R820T-master.zip R820T2 Breakout Board

- Elonics-E4000-Low-Power-CMOS-Multi-Band-Tunner-Datasheet.pdf e4000\_refsch\_rev4.pdf

In the old days rtlsdr sticks re-tuned relatively slowly. As time passed re-tuning speed has been increased by clean-ups in code and specifically keenerd's changes so the tuner doesn't wait nearly as long for the pll to settle. More recently tejez/s mod made the re-tuning even faster by updating all the changed registers for a re-tune in one r82xx\_write 12C call. With this done you can re-tune at rates of up to 41(1) hops per second; a ~2x improvement over then-existing drivers. Since then all of these re-tuning changes have been incorporated into the main rtlsdr.

Further massive speed-ups can be had at the cost of pretty much all reliability. By not waiting for PLL lock at all and always leaving the 12c repeater register enabled tejeez reports retuning speeds of up to 300 jumps per second are possible.

As of Aug. 2014 a handful of people have found ways to extend the r820f frequency range as well. Initially thought to top out at 1700 MHZ the R820T driver has has re-written to tune from 22 to 1870(1) MHz. While efforts have been made to extend the lower range as well, with the PLL seeming to lock down to 8 MHz in some cases, this range turns out to be full of images and repeats of the higher frequency range. A later effort with the addition of driver tweaks to the RTL2832 downconverter pushed the low end down to ~15 MHz. The code is at the code is at the restriction of the re

After tejeez worked out the no-mod HF reception a couple people have noted that the tuners with fc0013 receive HF even better than the R820T board designs. So if you have one of those laying around you might want to try HF with it

The E4K has settings for LNA (-5..+25dB), mixer (4 or 12dB) and total of 6 IF gain stages with various gains allowing for 1dB steps between 3 and 57dB. The software only deals with LNA and mixer gain and not independently. IF gain can be set through the API.

R820T also has LNA, mixer and IF gain settings - the exact steps are not known. The numbers in the library code are through measuring the gain at a fixed frequency. That gave 0.33dB for the LNA, 0.16dB for the mixer and -4.7.40.8dB for the IF gain. The current library does not expose these settings through an API, only LNA and mixer are set through some algorithm. IF gain is set to a fixed value.

both gives more detail about the R820T step size,

The mixer gain step is 1dB (matches the empirical data passably, but not great) and the IF/VGA gain step is 3.5dB (matches mine basically dead-on). LNA gain step is not mentioned, all it says is "1111 - max, 0000 - min"

All of the dongles have significant frequency offsets from reality that can be measured and corrected at runtime. My ezcap with E4000 tuner has a frequency offset of about ~44 Khz or ~57 PPM from reality as determined by checking against a local 751 Mhz LTE cell usin Scanner, Here's a plot of frequency offset in PPM over a week. The major component of variation in time is ambient temperature. With the R820T tuner dongle after correctly for I have has a ~ ~17 Khz offset at GSM frequencies or ~35 ppm absolute after applying a 50 ppm error correction. When using kalibrate for this the initial frequency error is often too large and the FCCH peak might be outside the sampled 200 KHz bandwidth. This requires passing an initial ppm error parameter (from LTE scanner) -e. Another tool for checking frequency error is determined by checking frequency error is often too large and the FCCH peak might be outside the sampled 200 KHz bandwidth. This requires passing an initial ppm error parameter (from LTE scanner) -e. Another tool for checking frequency error is often too large and the FCCH peak might be outsided the sampled 200 KHz bandwidth. This requires passing an initial ppm error parameter (from LTE scanner) -e. Another tool for checking frequency error is often tool are provided to the sampled 200 KHz bandwidth. This requires passing an initial ppm error parameter (from LTE scanner) -e. Another tool for checking frequency error is often tool are provided to the provided that the provided the provided the provided the provided that the provided the provided that the provided that the provided the provided that the

very cool is the MIT Haystack people switching to rtlsdr dongles (pdf) for their SRT and VSRT telescope designs, Use of DVB-T RTL2832U dongle with Rafael R820T tuner (pdf). The first of these characterizes the drift of the R820T clock and gain over time as well as a

As of 2015 there are a number of SDR-enthusiast targeting dongles produced with temperature controlled oscillators (TXCO) that run at less than 1 PPM with no start-up drift.

I recently (06-15-2014) found out from prog (of SDR# and airspy) that there are actually two different versions of the R820T tuner. The normal one and the R820T2. The T2 has different intermediate frequency filters allowing for wider IF bandwidths and apparently slightly be sensitivity (a few dB lower noise floor?). For risked dongles this difference in IF filter bandwidth usually doesn't matter much since all of them are larger than the RTL2832U's debug/SDR mode bandwidth of ~3 MHz. But there are certain situations where a larger tuner bandwidth advantageous: such as when using lowerts IHE tuning mod. As of Sept. 2014 some of the new R820T2 have been showly up in Terrater "F4000 upgrade" model sticks. But don't count on it. I bought one from ebay seller "smallpartsbigdifference" which had a photo showing an R820T2 and it was just an R820T. Since ~2015 R820T2 have been showled the R820T2 Register Descriptions.

As of 2017-12-11 Rafael Micro has been sending out emails saying the R820T2 has been discontinued. Alternate versions of the series, not pin compatible, are the R836, R840, or R828D. I've already seen SDR-targeted dongles using the R828D

As of March 2018 it's uncertain how many non-defect bin R820T2 tuners Rafael Micro has left but rtlsdrblog has said that if very large bulk orders could be made Rafael Micro might be willing to produce more

In Feburary 2015 Leif sm5bsz (of linrad) relased a modified librtlsdr with changes to the rtlsdr R820T tuner code to allow for finer grained control over IF filter settings.

The IF filter which actually is a low pass filter and a high pass filter can be set for a bandwidth of 300 kHz. Dynamic range increases by something like 30 dB for the second next channel 400 kHz away. It is also possible to get some more improvement by changing the gain

Following this gat3way's <u>patched gr-osmosdr</u> and Vasili <u>ru</u>'s <u>SDR# driver</u> were released. gat3way made the IF filter width variable from within gqrx by presenting it as a gain value. Vasili's rtlsdr SDR# driver also moves the SDR# decimation normally applied during demodulation to the front of the IQ stream. This gives better dynamic range for the visual FFT but demodulated quality is not changed. So far this is all experimental but expect it to be brought mainline on both sides soon.

keenerd's experimental branch automatically set IF filter width based on sample rate but had not exposed them as manually set values.

TLeconte.github.io made some great plots of the R820T2 IF filter shapes in, "Playing with the Airspy R820T IF bandwidth". It should apply just the same for rtlsdr with changes.

# R828D varian

In late 2013 Astrometra DVB-T2 dongles with the R828D tuner (pic) paired RTL2832U have begun to appear (2). The DVB-T2 stuff is done by a separate Panasonic chip on the same 12C bus. merbanan wrote a set of patches, <u>rtl-astrometa</u>, for librtIsdr has better support these tuners. The performance hasn't been characterized but it at least works for broadcast wide FM via SDR, stevelpn's preliminary testing suggests bad performance in the form of the crystal for the DVB-T2 demodulator leaking fixed spurs 25 dB above noise floor in the IF at approximately 196 and -820 KHz. He was able to mitigate these with the hardware mod of removing the crystal for the DVB-T2 chinical support was added to the rtl-sdr on Nov 5th while testing support was added on Nov 4th. In April 2017 u/strangerwithadvice on reddit made a quality post on the r/rtlsdr subreddit where he characterized the noise floor of an R828D dongle stock and with a number of modifications to reduce noise.

randomsdr reported on Freenode ##rilsdr IRC on 2015-09-03 that the Leadtek Winfast DTV2000DS PLUS pci card has 2x FC0013 tuners and 2x rtl2832u chips like 2 normal rtlsdr dongles. Performance is not good but tools like rtl\_fm work if the VID/PID is added to the rtlsdr driver table and udev rules set. It isn't recommended except as a novelty.

2012-08-22: The E4000 tuner datasheet has been released into the wild. Elonics-E4000-Low-Power-CMOS-Multi-Band-Tunner-Datasheet.pdf, but.

"All the ones that are documented in the DS are 'explained'in the driver header file ... And the rest, the datasheet call them "Ctrl2: Write 0x20 there" and no more details"

2012-09-07: Experimental support for dongles with the Rafiel Micro R820T tuner that started appearing in May has been added to rtl-sdr source base by stevem. These tuners cover 24 MHz to 1766 MHz. They also don't have the DC spike caused by the I/Q imbalance since they use a different, non-zero, IF. On the other hand, they might have image aliasing due to being superheterodine receivers. See stevens tuner comparisons. On 2012-09-20 the R820T datasheet was leaked to the ultra-cheap-sdr mailing list. The R820T Z Register Description pdf was provided by lugit areanges in the ultra-cheap-sdr mailing list after be received in from Rafaelhkincor. The officiar large is 42-1002 with a 3.5 dB noise figure. On 2012-10-04 my order arrived. The lifting this tuner every much since it actually works well, locking down to 15 Alba or so without direct sampling mode. Here's a rough gruppf spectral map of 24 to 1700 MHz over 3 days I made with some custom perl and python scripts. Don't judge the r820t on the quality of that graph, it is just to show the range. You can see what I think is either front-end mixer filters not attentionating enough or actual intermodulation as RFI. I do almost no processing of the signal (ie, no IQ correction), don't clear the buffer between samples (LSB probably bad), and use a hacky way to display timeseries data in gluplot. Real SDR software like SDR# shows them to be equal or better in quality to E4ks.

stevem did gain measurement tests with a few dongles using some equipment he had to transmit a GSM FCCH peak, "which is a pure tone." This includes the E4000 and R820T tuners. In addition he measured the mixer, IF and LNA for the R820T.

High Frequency (0-30Mhz) Direct Sampling Mod



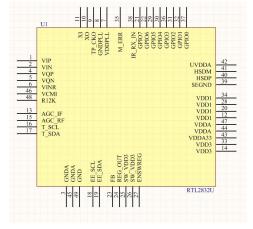
Steve Markgraf of Osmocom has created an experimental software and hardware modification to receive 0-30Mhz(\*) by using the 28.8 MHz RTL2832U ADCs for RF sampling and aliasing to do the conversion. In practice you only get DC-14.4MHz in the first Nyquist zone but the upper could be had by using a 14.4 MHz to 28.8 MHz bandpass filter. In the stereotypical ezcap boards you can test this by connecting an appropriately long wire antenna to the right side of capacitor 17 (on EzTV668 1.1, at least) that goes to pin 1 of the RTL2832U. That's the one by the dot on the chip surface. Apparently even pressing a wet finger onto the capacitor can pick up strong AM stations. This bypasses static protection among other things so there's a chance of destroying your dongle. For gr-osmosdr the parameter direct\_samp=1 or direct\_samp=2 gives you the two 1 or two Q inputs.

It has recently become possible to use direct sampling with no hardware modifications at all. It is still very experimental and performance is bad. In Oct 2012 Anonofish on the ririsdr subreddit had discovered the PLL would lock for a small ~ 3686.6 MHz - 3730 MHz range far outside the normal tuning range and there seemed to be signals there. In January 2014 ##rtlsdr BK Channel user tejeez figured out this bypassed the tuner (mixer leakage) and implemented a set of register settings (R820T IF frequency, IF filter bandwidths, r82xx\_write\_reg\_mask(priv, 0x12, val, 0x08) replaced with r82xx\_write\_reg\_mask(priv, 0x12, val, 0x10, 0x18)) that would exploit this to enable HF reception. Shortly thereafter keenerd assembled everything into a relatively easy to use patch-set.

If you want to give HF listening a try with no risk keenerd has added these changes rtl. fm and rtl. power in his experimental rtlsdr repository. To use the no mode mode with rtl\_tools append the argument, "-E no-mod". To use the no-mod direct sampling in something that uses grosmosdr, like gqrx or GRC flowgraphs, add the following to the the "device string" parameters: ie "direct\_samp=3". Plug your HF antenna into the normal connector, no hardware mods needed.



I've been told my pin numbering doesn't correspond to the datasheets, so take that with salt. The relative positions are correct regardless of the numbering. The RTL2832 ADC differential input impedance is ~3,300 Ohm



A number of people have tried to match the ADC's input impedance and both differential inputs by using baluns of various sorts. The datasheet seemed to say 200 Ohm so a lot of people (myself included) tried 4:1 baluns which did improve performance. But better matches can be found using the 1800 Ohm (ref: GENU).

Dekar has a page showing how to use an ADSL transformer to generate signal for the ADCs differential input using pin 1 (+1) and 2(-1) on the RTL2832. miking has a useful pdf schematic with part numbers for using wide band transformers or toroids for winding your own. Here's a series of posts from bhSea20th showing how to use a FT37-43 ferrite core. And another example from IW60VD Fernando. Py47BZ as well. The ADC has a differential/balanced input so this is done mainly for the unbalanced-balanced conversion. But the ADC input pins also have a DC offset so you can't just connect one to GND for that.

Tom Berger (K1TRB) used multiple core materials with trifilar wire and performed tests using his N2PK virtual network analyzer on May 19th (2013).

Hams love type 43 ferrite, but for almost every application, there is a better choice. For broadband HF transformers Steward 35T is generally a better choice. Therefore, I wound a couple transformers and did the comparison. Type 43 and 35T Transformer Material Comparison.

For my tests with direct sampling mode I ordered a couple wideband transformers from coilcraft. The PWB-2-ALB and PWB-4-ALB to be specific. I sampled the PWB-4-ALB for free and ordered 4 of the PWB-2-ALB for ~\$10\$ shipped. Both seem to work fine though I have no

If you're particularly interested in HF work then an upconverter would be better than the HF mod. With the mod there will be aliases(\*) for any frequency over 14.4 Mhz (1/2 the 28.8 clock rate). So you'd want a 14 MHz lowpass for the low end or a 14-28 MHz bandpass for the high end. And probably other little idiosyncracies. A lot of people chose to just use an upconverter instead. KF7LE wrote up short summaries comparing 16 popular upconverters.

Another alternative is to make a diplexer so that you get both HF via direct sampling and VHF+/etc without any switches. G8JNJ has a detailed guide with annotated photos on how to build the appropriate circuit and modify the latest R820T2 type dongles with it. He reports being able to receive from 15 KHz to 1.8 GHz with this mod.

# Removing the tuner entirely

Since the tuner is not used in direct sampling it can be removed entirely (especially in a case where it is dead). Goatman contacted me on ##rtlsdr IRC about the process and pin jumpers needed to feed the rtl2832u the clock directly from the onboard oscillator in his case.

tman> with a simple wire from pin 8 to 10 on the tuner pad tman> rtl\_test reports all is well with direct sampling tman> ... Gayx works fine btw with direct sampling even if no tuner is found



Using External Clocks and coherent sampling in general...

"I recently came up with a trivial hack to build a receiver with multiple coherent channels using the RTL dongles. I do this basically by unsoldering the quartz clock on the slave units and cable the clock from the master RTL dongle to the input of the buffer amplifier (Xtal\_in) in the slave units (I've attached some pictures)."

Since I've seen a lot of people asking, the dongles he used were Newsky TV28T v2 w/R820T tuners.

- \$16 dual-channel coherent digital receiver
   Passive radar with \$16 dual coherent channel rtlsdr dongle receiver

Max Manning also implemented a passive radar system using clock sharing rdsdr. Ben Silverwood also did so as, "Low cost RTL-SDR passive multistatic DAB radar," an implementation in matlab. The youtube video description has links to photos of the setup

Also, there's a Japanese seller with high precision SMD 28.8 MHz crystals. And an ebay seller with high precision 28.8 MHz oscillators for around ~\$30 shipped.

Things again became exciting in June of 2014. Going beyond simple clock sharing and if's max of 3 dongles, YO3IIU put up a great post his build of a 4+ dongle RTL2832u based coherent multichannel receiver using a CDCLVC1310-EVM dev board from TI for clock distribution. His post shows the results of a gnuradio block he coded that does all the correlation math to align the samples from each receiver (which are out of step due to the way USB works). Unfortunately the software was never released.

stevejm's experiments were the first I heard about back in 2011. He used his 13MHz cell-phone clock as a reference for a PLL to generate 28.8MHz. He said he used 1v peak to peak. He also related it was possible to not even use the PLL and just the 13 MHz clock if w/E4000 tuners if you don't care about sample rate offset.

<steve|m> not really, just a picture and a short clip: http://steve-m.de/projects/rtl-sdr/osmocom\_clocksource.webm\_http://steve-m.de/pictures/rtlsdr\_external\_clock.jpg
<steve|m> a motorola Cl39

The Green Bay Public Packet Radio guys have written up an interesting article on using 14.4 MHz temperature controlled crystal oscillators sent through a passive (two diode) frequency doubler followed by crystal filters made out of the old rtlsdr clock crystals to provide a low PPM error clock for rtlsdr devices. Since their mirror was missing images I cut them out of the Zine pdf and made a mirror here.

I first heard about the GBPPR article from patchvonbraun who implemented one and performed tests which he posted about on the Society for Amateur Radio Astronomy list. It turns out that even with a good distributed clock the 2x R820t rtlsdr dongles still have large phase error for some reason, see: Phase-coherence experiments with RTLSDR dongles and the photo post: Progress towards using RTLSDR dongles for interferometry.

Alex Paha has also done clock distribution but unlike the others he used E4000 tuner based receivers for his dual coherent receiver. He also seems to be using only half the I/Q pairs. This post is in Russian.

### Actually maintaining coherence over re-tunes and USB2 latency

### rtl\_coherent

In October 2015 teejez uploaded his rtl coherent code for maintaining multi-dongle coherence using external antenna switches to disconnect the antennas and connect all to a common noise source for correlation calibration. Here's a video of him using it to make a 3 dongle direction finder.

Each dither-disabled rtl-sdr is fed from the same reference clock. They still have unknown phase shifts and sampling time differences relative to each other. This is calibrated by disconnecting them from antennas and connecting every receiver to the same noise source. Cross correlation of the noise gives their time and phase differences so that it can be corrected. Currently the signal is received and processed in short blocks with each block starting with a burst of calibration noise.

As I understand it the switch chips are sa630 that "look" for dongle i2c traffic. There are controlled by two RC delay circuits so that every time you change frequency (causing i2c traffic) it disconnects antennas, waits for some time, feeds a pulse (just one edge from the logic chip) into all dongles, waits a bit more and connects the antennas back. You can see the evolution of his setup from this earlier prototype to this later prototype and finally the version used in his direction finder.

Every time you tune any two (or more) dongles to a new frequency there will be a tiny difference in the frequency each actually tuned to. The offset must corrected before trying to correlate them. If you don't it'll look like there's a constantly varying phase shift. Also don't forget to let the dongles warm up to equilibrium otherwise this additional temperature related frequency shift will cause changes even larger than relative tuning offset and you'll get the "random" phase shift again.

### Multi-RT

As of 2016 Piotr Krysik's "Multi-RTL" (github) has made maintaining coherence of multiple dongles accessible even to the amateur. His GNU Radio block handles all the complex details of keeping multiple rtlsdr coherent even when they're tuned to different frequencies and over returnes. It requires no external circuitry. You just have to distribute the clock signal with cable.

### PLL Dithering and you

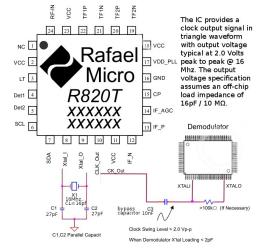
On the clock coherencey side Michele Bavano's has explored, tweaked, and replaced, librtlsdr's pll setting code, intermediate frequency, and PLL dithering settings, such that the math, and results, work out cleaner. Using this modified driver he was able to minimize frequency setting errors and improve his GPS carrier following code. This is written up with code examples at his blog in, GNSS carrier phase, RTLSDR, and fractional PLLs (the necessary evil). Without dithering you can only tune to increments of 439.45 Hz. With dithering, you can tune to approximately anything.

tejeez from the ##rtlsdr IRC relates that this can be done in r82xx\_set\_pll by changing r82xx\_write\_reg\_mask(priv, 0x12, val, 0x08) to r82xx\_write\_reg\_mask(priv, 0x12, val, 0x10, 0x18). This has been implemented as an option in rtl\_sdr, '-N', in keenerd's experimental branch.

### Misc

In the absence of any useful information about the RTL2832U clock here's some information about the R820T's clock system.

Crystal parallel capacitors are recommended when a default crystal frequency of 16 MHz is implemented. Please contact Rafael Micro application engineering for crystal parallel capacitors using other crystal frequencies. For cost sensitive project, the R820T can share crystal with backend demodulators or baseband ICs to reduce component count. The recommended reference design for crystal loading capacitors and share crystal is shown as below.



# Noise, shielding, cables, and why is that FM signal there?!

When you see something weird, like commercial FM broadcasts at 27 MHz, what you are seeing incomplete filtering of mixing products. It's the harmonics of the square wave driving the mixers combined with insufficient rf filtering to suppress the response. You can tell if it is a local oscillator mixer harmonic leakage by sweeping the frequency and seeing how fast the ghost signal moves relative to this; look for linear relationships (ie, 2x the speed, 1/4 the speed, in-depth reference). Sometimes local signals can be powerful (ie, pagers) or close enough to make the preamplifier behave non-linearly resulting in intermodulation. For this kind of RFI turning down the gain helps.

The tuners all have a certain amount of intrinsic noise too. keenerd had done tests with an R820T rtlsdr terminated to a resistor inside of a metal box. For these tests rtl\_power gain was set to max (49.6dB) and a frequency sweep was done through the entire tuner range, 1820t Background Noise. The 28.8 MHz spikes from the clock frequency can be seen among other abberrations.

But not everything is a ghost from hardware design problems. Depending on your computer setup and local electronics there could be a lot of "real noise"; LCD monitors are a common culprit for VHF noise spikes distributed across wide ranges. It is best to shield and put ferrites on everything if you can

To solve the commercial FM mixing problems an FM trap can be used. Commercial ones work fine typically. But for non-commercial FM RFI like emergency services and pagers custom filters must be made or ordered. Adam-9A4QV has a detailed write-up on making FM trap with a very high upper passband (all the way to 1.7 GHz) with links to design for other low VHF bands. tejecz shared his VHF bandstop design on IRC. Like Adam's it has the unique feature of not also wiping out harmonics of the FM band: fm-notch\_ipg fm-notch\_schematic.png. This means you can use it and still do wideband frequency hopping (unlike, say, a 1/4th wave coaxial stub). For more information on this general type of coaxial cable notch filter check out Ed Loranger's write up on VHF Notch filters (photo). For my powerful 461 MHz RFI that can be received without an antennal 1 use a custom 3 cavity notch filter from Par Electronics.

Actinonyx describes one way to doing this using a single strip of aluminum tape combined with a spring to connect it to the dongle ground. Also Czermann at the sdrformariners blog made a somewhat confusing but definitely empirical comparison of noise levels compared to different hardware mods like disconnecting the USB ground from the rulsdr ground. Quite a few people have had success with that and scotch tape around the USB connector works to test it.

Some others bond the enclosure to both the antenna and the USB shield and this works reliably and well.

Martin from ggininet finds the most effective mod to reduce USB and DC-converter noise is shielding the antenna input area with metal soldered to the pcb ground, "The noise seems to be coupled directly between components on the topside of the PCB." You can find it if you scroll about halfway down the page linked.

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### 8.4. RTL2832U Internal Switching Regulator

The RTL2832U integrates a switching regulator with input voltage 3.3V to output voltage 1.2V. Figure 6 shows the application circuit.

The ENSWREG pin default power is 3.3V. Applying 0V turns off the switching regulator.

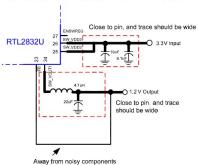


Figure 6. Internal Switching Regulator Layo

Additional noise comes from the switching power supply in the RTL2832U that runs at 1.024MHz. This drops the supplied 3.3v down to the 1.2v needed for internal use. <u>ttrflech has successfully disconnected this switching supply</u> replaced it with 3 diodes to drop the 5v line down to ~1.2v. In the example linked above ttrflech uses power form the far side of the board but the eeprom's power rail would also work. This decreases spurs in HF significantly. It will increase power usage though; something to watch out for when R820T dongles start out at ~300mA a

Laidukas's "Mods and performance of R82012 based RTL SDR receiver covers replacing all the power rails with external linear regulators, increasing the amount of bypass capacitance on power lines, adding extra chip filtering for the USB 5v line, cutting off the IR receiver part of the PCB, wiring in a TCXO 28.8MHz oscillator, creating a shield with kapton tape and copper foil soldered extensively to the PCB ground, and a new heavy metal case and connectors.

To reduce signal loss over long distances and get away from computer RFI like to run long USB \*active\* extension cable\* with hubs at the end and ferrites added instead of coaxial cable. Around this USB cable I clip on 5 or 6 ferrites at each end. Active extension/repeater USB2 cables of up to 25m in length can be used.



When I want to do some scanning that takes advantage of the tuner's very wide ranges I use five types of antenna: discone, spiral, dual planar disks, vivaldi (tapered slot), and horns (TEM and pyramidal). Discone, dual planar disk, and archimedian spiral antenna can omnidirectionally cover almost the full range of the E4000 tuner but things get a bit too large to go all the way to the 24 Mhz of the R820T. You can refer to the seperate spiral antenna page for construction and technical details. To build my discone of I followed Roklobsta's D.I.Y. Discone for RTLSDR (archive or grammor). With just a discone and rIL power it's possible to see lots of LEG satellite carrier frequencies doppler across the spectrum.



To get an idea of how much you can see with a discone here's a directory where I produce -2 to 4 day long -70 to 1000 MHz range 25KHz resolution 45k\*10k pixel spectrograms. They each have a javascript zoomable interface to load small tiles progressively. An example. With just a discone and rtl\_power it's possible to see lots of LEO satellite carrier frequencies doppler across the spectrum.

But with a band specific helix in a cone reflector (helicone) many more satellites can be picked up. The previous is a link to a zoomable spectrogram of ~2 days of the 1616-1626 MHz satellite band that Iridium satellites use. No LNA was used. There's plenty of RFI/EMI even through a 1 GHz high pass but the satellite doppler passes are clearly there in numbers if you zoom in far enough.

When using such broadband antenna, or even a band specific helix, it is possible to pick up powerful out of band signals due to overloading or incomplete mixer filtering. It's important to identify any extraordinarily powerful transmitters nearbye and filter them out. In my case I have a 50w transmitter at 461 MHz across the street always going full power. I bought a custom tuned 3 cavity notch filter from PAR Electronics. This limits the upper frequency range to 1GHz but does at least solve the RFI problem.

Usually the spectra are much cleaner when using directional and resonant antenna instead of wideband omnidirectionals. But many directional antenna like helix and log periodic dipoles have very large out of band sidebands on low frequencies not in the designed range

# Page Sections

- Pagers
  Ggrx on Ubuntu 10.04
  LTE Scanner on Ubuntu 10.04
- LTE Scanner on Ubuntu <u>Dongle Logger</u> GNU Radio Companion
- Clocks Broadband Antenna

# RTL-SDR Links

- <u>rtl-sdr wiki</u> osmocom <u>Known applications rtl-sdr wiki</u> osmocom

- Known applications trustory with riskforg with introduced to riskforg with processing the riskforg region of the riskforg region of the riskforg risk

Warning: I'm learning as I go along. There are errors. Refer to the proper documentation and original sources first.

# GNU Radio \*and\* RTL-SDR Setup

You don't need GNU Radio to use the rtlsdr dongles in sdr mode, but there are many useful apps that depend on it, patchyonbraun has made setting up and compiling GNU Radio and RTLSDR with all the right options very simple on Ubuntu and Fedora. It automates grabbing the latest of everything from git and compiling, It will also uninstall any packages providing GNU Radio already installed first. Simply no, http://www.shrac.org/files/build-guuradio, and irll automate downloading and compiling of prequisites, libraries, correct git branches, udev settings and more. I had no problems using Ubuntu 10.04, 12.04, or 14.04. These days (2015) pybombs is slowly taking over for build-guuradio but for now this works best.

If you're thinking about trying this in a virtual machine: don't. If you do get it partially working it'll still suck

As an aside: If you're an OSX user then you can use the MacPorts version of GNU Radio (including gqrx, etc) maintained by Michael Dickens.

# default is latest 3.7 # install old 3.6.5.1

Install 3.7. Most gnu radio projects have been ported to it as default. Only a few old things will require 3.6.

An (rg)install looks like this. It might be useful to save the log output for future reference. Then test it. The test output below is from a very old version of rtl\_test with an E4K dongle. Newer versions, and R820T tuners will output slightly different text.

rtl\_test -t Found 1 device(s): 0: ezcap USB 2.0 DVB-T/DAB/FM dongle Using device 0: ezcap USB 2.0 DVB-T/DAB/FM dongle Found Elonics E4000 tuner

Once GNU Radio is installed the "Known Apps" list at the rtl-sdr wiki is a good place to start. Try running a third party receiver, a python file or start up GNU Radio Companion (gnuradio-companion) and load the GRC flowcharts. If you're having "Failed to open rtlsdr device #0"

Стр. 5 из 21 20.11.2020, 23:11 errors make sure something like /etc/udev/rules.d/<u>15-rtl-sdr.rules</u> exists and you've rebooted.

When updating you can just repeat the install instructions which is simple but long. The advantage to repeating the full process is mainly if there are major changes in the gr-osmosdr as well as rtl-sdr. It'll do things like Idconfig for you

If you don't have the patience for a full recompile and there haven't been major gnu radio or gr-osmosdr changes it's much faster just to recompile rtl-sdr by itself. The instructions to do so are at the osmosdr page. It'll only take a few minutes even on slow machines. Once you have the latest oir clone it is like most cmake projects:

git clone git://git.osmocom.org/rtl-sdr.git cd rtl-sdr; mkdir build; cd build; cmake ../; make; sudo make install; sudo ldconfig

### rtl-sdr supporting receivers, associated tools

### keenerd's rtl-sdr branch

This experimental branch contains a number of useful low processing power utilities, expansions of the original rtl tools, and improvements to the R820T driver re-tuning speed. A lot of them have already been merged into the librtlsdr master but rtl\_fm and rtl\_power fixes, features and bugs appear here first. If fm is for scanning, listening, and decoding (and not just FM), rtl ados for plane watching with an external ads-b viewer, rtl ceprom for checking and "setting" serial numbers and related data if you dougle has an eeprom. And as of 2013-09-20, rtl, power, a total power frequency scanner. The total variables tools are very good for slow machines or when you want to do command line automation. Just build it like the <u>commoner rtlsdr page does</u> for the vanilla install. Use these on the rapberry pi.

git clone https://github.com/keenerd/rtl-sdr.git git clone https://github.com/keenerd/rtl-sdr-misc

cd rtl-sdr; mkdir build; cd build; cmake ../ -DINSTALL\_UDEV\_RULES=ON; make; sudo make install-udev-rules; sudo ldconfig

Most people use rtl\_power for smaller total bandwidths (<200 MHz) and higher spectral resolution using the default FFT mode. This is visualized with keenerd's heatmap.py and can result in some really impressive plots when done with ~25% crop mode. Just refer to the -h "help" in rtl\_power for instruction. There is also an rtl\_power guide at keenerd's website.

For RMS average power mode, which kicks in automatically for FFT bin sizes 1 MHz and larger, 1 do visualization of the resulting .csv file with gnuplot. Because the entire bandwidth is summed and saved as one value the the data rate to disk, and spectrogram dimensions are much lower than FFT mode.

Example gnuplot visualization, annotated, and the gnuplot format, and colour palettes used to generate them.

If you do a large number of frequency hops, (hundreds) then the time adds up. On my two computers the R820T tuner dongles average about 55 milliseconds per retune and sample cycle. I sometimes have dongles that'll fail to lock pll and go into a loop. The -e parameter sets a time limit for a run. Combining this time limit with a bash while loop results in pretty low downtime with resiliance to rtlsdr and USB failures.

# 0.055 seconds \*((1724-24) MHz/(1 MHz)) = 93.5 seconds = "-i".
while true; do ./rtl\_power -f 24M:1724M:1M -i 94 -g 20.7 -p 30 -e 1h >> 2013-12-02\_totalpower\_r820t\_24-1724\_1M\_94s.csv; done

To combine the results from multiple dongles just cat the files together. But on gnuplots end each new .csv filename requires you to manually edit the gnuplot format. Additionally you need to set the output spectrogram filename and a pixel width. I find for 1000 Mhz @ 1 MHz that approximately 1000px per 100 MB of file size is required to cover all gaps.

And that pops out a png.

For rtl fm stuff refer to keenerd's site's Rtl fm Guide

• Spektrum: an rtl\_power GUI frontend:
• There are a lot of rtl\_power GUI frontends but the most useful for me is Spektrum. It uses a modified rtl\_power with a Processing GUI front-end. It's available for limux and windows. One of it's best features is the "relative mode" for use in measuring changes in antennas

ivonbraum (Marcus Leech)'s multimode:

AM, FM, USB, LSB, WFM, TV-FM, PAL-FM. Very nice, easy to use (screenshots: main, scanning). It has an automated scanning and spectral zoom features with callbacks to click on the spectrogram or panorama to tune to the frequency of interest. There's a toggle for active gain control too. The way to get it is,

then instead of using GRC, just run the multimode.py as is.

If you run it outside of the svn created directory you might need to append ~/bin to pythonpath to find the helper script. If you used build-gnuradio it'll tell you what this is at the end of the install.

Alternately set it in your ~/.bashrc. If you do the below make sure to reload in the terminal by "source ~/.bashrc")

PYTHONPATH=/usr/local/lib/python2.6/dist-packages:~/bin export PYTHONPATH;

When setting the sample rate it is rounded-down to a multiple of 200 Ksps so the decimation math works out

python multimode.py --srate=2.4M # use normal mode with 2.4 MHz bandwid ./multimode.py --devinfo="rtl=0,direct\_samp=1" # use direct sample mode

If you have overruns like "OOOOoo..." then try reducing the sample rate or pausing the waterfall or spectrum displays

"The audio subsystem uses 'a' as the identifier, and UHD uses 'u'. With RTLSDR, it'll issue 'O' when it experiences an overrun. Which means that your machine isn't keeping up with the data stream. Sometimes buffering helps, but only if your machine is right on the edge of working properly. If it really can't, on average "keep up", no amount of buffering will help."

If you have overruns like "aUaUaUa" or just "aaa" then the audio system is asking for samples at a higher rate than the DSP flow can provide (44vs48Khz, etc). Use "aplay -I" to get a list of the devices on your system.

The hw:X.Y comes from this mapping of your hardware -- in this case, X is the card number, while Y is the device number. Or you can use "pulse" for pulseaudio. Try specifying,

• ggtx:

• Written by <u>Alexandru Csete OZ9AEC</u> "gqrx is an experimental AM, FM and SSB software defined receiver". The original version did not have librtlsdr support so changes were made by a number of others to add it. A couple weeks later Csete added gr-osmosdr support to the original. Dekar established a <u>non-pulseaudio port of ggrx for Mac OSX</u>, GNU Radio 3.7 has recently been released and it is not exactly backwards compatible, patchvonbraun's build-gnuradio.sh pulls 3.6.5 3.7.x by default.

As of August 9th 2013 Garx 2.2.0 has been released. This upgraded version can now be installed as binaries with all of it's dependencies pre-packaged on both Ubuntu linux (a custom PPA, no 10.04 packages) and Mac OS X! That includes all the GNU Radio stuff. So this is an all-in-one alternative to building GNU Radio from source.

I think this person's guide is better than mine.

git clone https://github.com/csete/ggrx.git cd ggrx

# on Ubuntu/Debian. 8udo cet .... cd gqrx

f on Ubuntu/Debian, sudo apt-get install qtcreator , if you don't have it.
qtcreator gqrx.pro f press the build button (the hammer)
f or avoid qtcreator and do it manually. if you have qt5 too, use "qmake-qt4"
mmake

# rtlsdr w/Gqrx on N900 phones xes provides pre-compiled packages of Gqrx and the GNU Radio dependencies for N900 linux cell phones

• SDR#:

• Written by prog (Youssef) for Windows. It is probably the best general purpose software for rtlsdr devices. Mono is slow and ugly on linux but if you restrict the sample rate it works fine. It's probably the easiest program to use, has the most diverse plugin ecosystem (example: Yasili\_ru's plugins), and has the best DSP and features for dealing with the quirks of the rtlsdr dongles.

As of 2015-09-14 the changes to Mono 4 allow SDR# to be viable to run on linux again. Make sure you have the latest Mono 4 though. This still requires soft linking in your system rtlsdr and portaudio library to the sdrshape.exe dir like below,

ln -s /usr/local/lib/librtlsdr.so librtlsdr.dll
ln -s /usr/lib/x86 64-linux-gnu/libportaudio.so.2 libportaudio.so

Just make sure you link your actual system rtlsdr and libportaudio, not my example path above. On debian/ubuntu find it by using locate,

Roklobsta's rtlsdr.org has a more detailed SDR# linux configuration guide if you're having trouble.

• Update: As of 2015-10-15 ADBS# is no more, ADBS# is no more, ADBS# is another easy to use application by prog, but specifically for plotting aviation transponders like gr-air-modes does. The distributed binaries also runs under linux with mono (or native in windows) and output virtualradar compatible data on 127.0.0.1:47806. If your antenna condition is crappy, try using filter = 1.

ger fosphor:
 0 ger fosphor is an amazingly fast and information dense spectrogram and waterfall visualization using OpenCL hardware acceleration. It surpasses the Wx widget elements in performance, and so usability, by far. With this visualization you can easily skip through 1 GHz of spectrum very quickly and actually notice transient signals as they pass. Right now it is not very configurable, just arrow keys for scale. But expect this to be the preferred visualization block in the future. I have written up an <a href="https://documents.org/lane/burger/barbors/burger/barbors/burger/barbors/bar

• A decoder of aviation transponder Mode S including ads-b reports near 1090 Mhz. It can be coupled to software to show plane positions in near real time (ex: VirtualRadar). This works under mono on Ubuntu 12.04 but not 10.04. Originally written by Nick Foster (bistromath) and adapted to rtlsdr devices first by Steve Markgraf (stevem), bistromath later added rtlsdr support. Here's an example of basic decoding done with the stock antenna on the early version by stevem. Nowdays it's better to use bistromaths'.

As of July 23, 2013 there was a major update to gr-air-modes which now includes a nice google maps overlay and works on gnu radio 3.7 branch only

git clone https://github.com/bistromath/gr-air-modes.git

Here's an example of install process and first run looks like.

# -d stands for rtlsdr dongle, location is "North,East" uhd modes.py -d --location "45,-90"

To use with virtual radar output add the below -P switch. Then open up virtualradar with mono and go to tools->options->basestation and put in the IP of the computer running uhd\_modes. There are not many compatible planes in the USA so far so even if you are seeing lots of Mode-S broadcast in uhd\_modes you might not see anything in virtualradar. Sometimes my server is running at <a href="superkuh.com:81/VirtualRadar/GoogleMap.htm">superkuh.com:81/VirtualRadar/GoogleMap.htm</a>.

uhd\_modes.py -d --location "45,-90" -P mono VirtualRadar.exe

o Dump 1090 is a Mode S decoder specifically designed for RTLSDR devices.

Antirezs' ADS-B program is really slick. It does not depend on GNU Radio, has a number of interactive modes, and it even optionally runs it's own HTTP server with googlemaps overlay of discovered planes; no virtual radar needed. It uses very little CPU and has

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impressive error correction. This is your best choice to play with plane tracking quickly.

```
# run cli interactively and create a googlemaps server on localhost port 8080
//dump1090 --aggressive --interactive --net-http-port 8080
# submit plane data to a tracking server
//dump1090 --aggressive --raw | socat -u - TCP4:sdrsharp.com:47806
```

· A guide on how to use linrad with rtl-sdr, and a modification to the the librtlsdr e4000 tuner code to disable digital active gain! reddit thread

"I tried various bits blindly and found a setting that eliminates the AGC in the RTL2832 chip. That is a significant part of the performance improvement."

As of 2012-07-07 this feature was added to the main (librtlsdr) driver as well.

o "This is an LTE cell searcher that scans a set of downlink frequencies and reports any LTE cells that were identified. A cell is considered identified if the MIB can be decoded and passes the CRC check."

"LTE-Tracker is a program that continuously searchers for LTE cells on a particular frequency and then tracks, in realtime, all found cells. With the addition of a GPS receiver, this program can be used to obtain basic cellular coverage maps

The author had only tested it on Ubuntu 12.04 but with some frustrating work replacing cmake files and compiling dependencies I made it work on 10.04. Scanner is very useful to get your dongle's frequency offset reliably and Tracker is very pretty. Remember to let your rtlsdr dongle warm up to equilibrium temperature before checking frequency error.

o "Kalibrate, or kal, can scan for GSM base stations in a given frequency band and can use those GSM base stations to calculate the local oscillator frequency offset."

The code was written by Joshua Lackey and made rtlsdr accessible by stevem. There is also a windows build made by Hoemchen. Let your rtlsdr dongle warm up to equilibrium temperature before running the test. When you're using this to find your frequency error it's important to use the -e option to specify intial error. 270k of bandwidth is used for GSM reception and if the error of the dongle is too large the FCCH-peak is outside the range. I compiled some install process and example usage notes.

Simple FM (Stereo) Receiver
 simple\_fim\_rev also by patchvonbraun is the best sounding and tuning commercial FM software in my opinion. He released a major update to his gnuradio creation at the end of October.

```
svn co https://www.cgran.org/svn/projects/simple_fm_rcv
cd simple_fm_rcv/
cd trunk
cd simple_w_-
cd trunk
less SEADMS
make install
## it'll install to ~bin/, so I use ~/superkuh/bin below
set PTHONEXTHF/usr/local/lib/python2.6/dist-packages:/home/superkuh/bin
# run the python script
# run the python script
for edit it ____cot.py
for edit it gouradio-companion simple_fm_rev.grc
```

### · my DongleLogger:

o I wrote these scripts do automatic generation of 1D spectrograms, per frequency time series plots of total power, and 2D spectral maps over arbitrary frequency ranges using multiple dongles at once. There is almost no DSP done and it is very simple but the wideband spectrograms and time series can be informative and fun regardless. It uses gnuplot for graphics generation. **Obsolete. Use rtl\_power instead.** 

You can see a typical gallery output at, http://erewhon.superkuh.com/gnuradio/live/ but more useful are the gnuplot spectrograms it can make.

o A simple, GRC-based tool for small-scale radio astronomy, providing both Total Power and Spectral modes. It has a graphical stripchart display, and a standard FFT display. It also records both total-power and spectral data using an external C program that records the data along with timestamps based on the Local Mean Sidereal Time.

This is another incredible tool by patchvonbraun. It does all the heavy lifting of integration over time and signal processing to get an accurate measurement of absolute power over a range. With it he has managed to pick out the transit of the milky way at the neutral hydrogen frequency using rtlsdr sticks and a pair of yagi antenna. The log file format is text and fairly easy to parse with gruplot but it comes with process simple tydar for cutting it into the bits you want and making total power or spectral component graphs. It'll make a directory called "simple\_ra\_data" in your home by default. Don't forget to set the --devid to rtl otherwise gnuradio won't find the gr-osmosdr source and it'll substitute a gaussian noise source.

```
git clone https://github.com/patchvonbraun/gr-ra_blocks cd gr-ra_blocks/
cmake . 7 make ; sudo make install ; sudo ldconfig
  cmake .; make ; sudo make install ; sudo ldconfig
cd ..
git clone https://github.com/patchvonbraun/simple_rs
cd simple_ra;
   make
make install
  # Using a single E4k tuner
//simple ra --devid rtl=0,offset_tune=1 --longitude -45
# Using a second dongle that just happens to have a R820T tuner
//simple ra --devid rtl=1 --longitude -45
# Use direct sample mode 8 1.25 MHz, log once per second, set longitude
# Use direct sample mode 8 1.25 MHz, log once per second, set longitude
//simple ra --devid rtl=0,direct samp=1 --freq 1.2566 --longitude -90 --lrate 1
# generate multi-day averaged qmplots using siderest itime
process simple type up: 00:00 0 2.0 - t "11 GHz Test" -f light.png -/simple_ra_data/tp-20130329-*.dat -/simple_ra_data/tp-20130328-*.dat -/simple_ra_data/tp-20130327-*.dat
//simple sa = 16 (README)
```

o Ear to Ear Oak made this wideband total power scanner that generates 1D spectrum plots over any tunable ranges with arbitrary integration times. It can update a matplotlib python plot GUI in real time and has the ability to output cvs values as well as an internal format. It's very useful for finding what's broadcasting in your area quickly. Using it's csv output and gnuplot I visualized a scan from \$54-1100\$ MHz.

If you want to use the data in gnuplot you have to sort it and make sure the header is commented out.

sort -n 54-1100\_500ms.csv > sorted\_54-1100\_500ms.csv

Here are some example gnuplot formats for the data

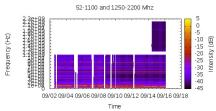
You can comment out the header manually but I instead prefixed a hash to the log writing behavior at line 786,

handle.write("# Frequency (MHz),Level (dB)\n")

# Pager stuff

• Thomas Sailer's multimon, "Linux Radio Transmission Decoder" which I use to (try to) decode pager transmissions around 930Mhz. And more recently <u>Dekar's multimonNG</u>, a fork with improved error correction, more supported modes, and \*nix/osx/windows support. Dekar also supplied a <u>GRC receiver for pagers to decode pager transmissions in real-time using fifos.</u> zarya has made <u>rtl. flex.py</u>, a gnuradio based flex decoder for pagers. It can be used to decode dutch p2000 messages, for example. This fills a gap in multimon-ng pager support.

# DongleLogger: my pyrtlsdr lib based spectrogram and signal strength log and plotter



Obsolete. Use rtl\_power instead. Automatic generation of and html gallery creation of wideband spectrograms using multiple rtlsdr dongles to divide up the spectrum. It also produces narrow band total charts, and other visualizations

# $(not\ live): \underline{http://erewhon.superkuh.com/gnuradio/live/} - click\ the\ spectrograms\ for\ time\ series\ plot$

These scripts cause the rtlsdr dongle to jump from frequency to frequency as fast as they can and take very rough total power measurement. This data is stored in human readable logs and later turned into wideband spectrograms by calling gnuplot. In order to further increase coverage of any given spectrum range multiple instances of the script can be run at once in the same directory adding to the same logs. Their combined output will be represented in the spectrogram.

# Details:

I don't know much python but the python wrapper for librtlsdr gyrtlsdr was a bit easier to work with than gnu radio when I wanted to do simple things without a need for precision or accuracy. Actualy receivers with processing could be made with it too, but not by me. This is the gist

power = 10 \* math.log10(scan[freq]) = scan = self.scan(sdr, freq) = capture = sdr.read samples(self.samples) = iq = self.packed bytes to iq(raw data) = raw data = self.read bytes(num bytes)

# The pyrtlsdr library can be downloaded by,

git clone https://github.com/roger-/pyrtlsdr.git

I have used the "lest.py," matplotlib graphical spectrogram generator that came with pyrtlsdr as a seed from which to conglomerate my own program for spectrum observation and logging. Since I am not very good with python I had to pull a lot of the logic out into a perl script. So everything is modular. As of now the python script generates the spectrogram pngs and records signal strength (and metadata) in frequency named logs. It is passed lots of arguments.

These arguments can be made however you want, but I wrote a perl script to automate it along with a few other useful things. It can generate a simple html gallery of the most recent full spectral map and spectrograms with each linked to the log of past signal levels. Or it can additionally generate gnuplot time series page (example) and link those intend of the raw logs. It also calls <u>TIF Cell Scanner</u> and parses out the frequency offset for passing to graphfreq.py for correction. I no longer have it running because of the processor usage spikes which intendally tasks. In the past I'd have reynou pudating the public mirror with a big pipe every "40 minutes."

As it is pyrtisdr does not have the get/set functions for frequency correction even if I sent the PPM correct from the perl script. Since the hooks (?) were already in libritsdr.py (line 60-66) but just not pythonized in rtlsdr.py they were easy to add to the library. These chang required to use frequency correction and make the int variable "err\_ppm" available. I have probably shown that I don't know anything about python with this description.

I forked roger's pyrtlsdr on github and added them there for review or use, <a href="https://github.com/superkuh/pyrtlsdr/commit/ffba3611ef0071dee7e1efee5c1a582e1e344c61">https://github.com/superkuh/pyrtlsdr/commit/ffba3611ef0071dee7e1efee5c1a582e1e344c61</a>, I apologize for cluttering up the pyrtlsdr namespace with such trivial changes but I'm new to this and github doesn't allow for private repositories.

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- rtl\_power was recently (2013-08-20) released by keenerd. It does most of what my scripts do, except much better, faster, and easier. I highly recommend you try it first.
   RTLSDR Scanner by Ear To Ear Oak is awesome for generating 1D wideband spectrograms.
   Enoch Zembecowicz made a polished and useful sdr logging tool
   Pantelities fri day sea can is another tool like RTLSDR Scanner for 1D total power scans. It is a good reference for using librtlsdr with C.
   If you are serious about measuring total power over one 2.5 Mhz range then simple\_ra, or simple radio astronomy, is best.

### fast version: see below

- · donglelogger-faster.tar.gz all needed files including pyrtlsdr
- radioscan\_faster.pl pyrltsdr using script; frequency setting and incrimenting, sampling, and logging.
   graphfrees\_faster5.py option passing, log parsing, plot making, frequency corrections wrapper, html image gallery generation of graphfrees\_gnuplot.py legacy functions

- graphfreqs.py pyrltsdr using script; sampling, and logging radioscan.pl manages graphfreqs, parses logs, makes plots, gets frequency corrections, generates gallery

### The faster version

### Speed ups, Inline C usb reset, and avoiding dongle reinitialization... (less options)

### cli switches/options

```
-dev 1 :: rtladr device ID (use to pick dongle)
-g 30 :: gain
-g 30 :: quantification of the provided of the p
```

These two scripts do fast scans within python from x to y frequency. Enabled it with -fast and make sure to set start and stop frequency with -f1 and -f2. Do not use -flist with this option

S while true; do perl radioscan faster.pl -d2 /tmp/faster -f1 25 -f2 1700 -fast -g 30 -r 2400000 -s 1.2 -m -p; sleep 1; done; Running graphfreq in non-batch fast mode 25 to 1700 Mm at 1.2 Mmz spacings.

Spectrograms and text output disabled.

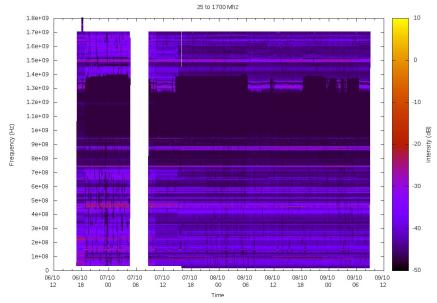
Using LTR Coll Scanner to find frequency Offset from 751 Mmz station...Found Rafael Micro R820T tuner

Generating Spectral Rap.

Generating Spectral Rap.

python ./graphfreqs\_faster.py 40000000 2400000 30 -58 /tmp/faster 1700000000 Found Rafael Micro R820T tuner ... (repeat many, many times)

This is an example output "spectral map" (a spectrogram with a silly name).



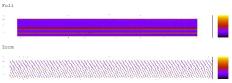
This example output above shows the overloading effects of using a wideband discone that picks up off-band noise. Each column is made up of small squares colored by intensity of the signal. Since the scripts start at the low frequency and sweep to high there is a small time delay between the bottom and top (see it more clearly zoomed in). And this is represented as the slant of the row. Sometimes strong signals will swamp out others resulting in discontinuities displaying as small dark vertical bands.

Or fast (-fast) scan a smaller range with smaller range (-f1,-f2; 24-80Mhz), with smaller samplerate (-r. 250 Khz) at smaller intervals (-s: 400Khz steps) with a gain of ~30. Only output a large spectrogram of all frequencies to the directory specified with -d2 as spectral-map.png. This example does not use frequency offset correct (-c) for even faster speeds.

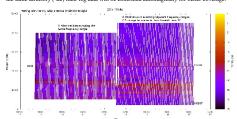
while true; do perl radioscan faster.pl -d2 /home/superkuh/radio/2012-02-02\_R820T\_Discone\_lowfreq -fl 24 -f2 80 -fast -g 30 -r 250000 -s 0.4 -m -p; sleep 1; done; Running graphfreq in non-batch fast mode 24 to 80 Mhz at 0.4 Mhz spacings. Spectrograms and text output disabled. Generating spectral map.

python ./graphfreqs\_faster5.py 24000000 250000 30 0 /home/superkwh/radio/2012-02-02\_R820T\_Discone\_lowfreq 80000000 400000 Found Rafael Micro R820T tumer Exact sample rate is: 250000.00014 Hz

# Combining multiple rtlsdr devices for greater speed



By splitting up the spectrum into multiple smaller slices and giving them to multiple dongles the time required for one scan pass can be greatly improved. The above spectrogram is made with 2 dongles, one for the lower half and one for the upper. It is from ryannathans who also contributed the code for for specifying device ID. This is as simple as running the script wice but giving each instance a different "-dev" argument to specify device ID. You can run as many rtlsdr devices with my scripts as you wish (up to the USB and CPU limits). If they are using the same directory (-d2) their log data will be combined automagically for better coverage.



while true; do perl radioscan\_faster.pl -d2 /home/superkuh/radio/2013-06-09\_multidongle -f1 25 -f2 525 -fast -g 40 -r 2400000 -s 1.2 -m -dev 0; sleep 1; done;

Found Rafael Micro R820T tuner

while true; do perl radioscan\_faster.pl -d2 /home/superkuh/radio/2013-06-09\_multidongle -fl 525 -f2 1025 1100 -fast -g 29 -r 2400000 -s 1.2 -m -dev 1; sleep 1; done;

Стр. 8 из 21 20.11.2020, 23:11 Sometimes I get corrupt samples that show a signal level of +60dB. These skew the scale of the output spectrograms. If I notice that they have occurred during a long run I'll use grep to find them and remove them manually. I replace the signal level with the level of the previous non-corrupt sample. In the future I'll build this kind of outlier removal in to the scripts, or sanity check before writing them.

```
grep -rinP " (3|4|5|6)\d+\.\d+" *.log
```

All the incremental improvements in speed I've made above are okay but not very easy to maintain with multiple script types (bash/perl/python). I'm slowly putting together an Inline C based perl wrapper for exposing librtlsdr's functions within a perl script to write this as a standalone in perl. This is slow work because I've never done anything like it before.

# rtlsdrperl - what if there were a perl wrapper for librtlsdr?

Well, there never will be. But here's some example code anyway.

### Older version

### graphfregs.py

You have to have the modified pyrtlsdr with the get/set functions for frequency correction. <u>ITE Cell Scanner</u> should also be installed so the "CellSearch" binary is available. Then download the two scripts above and put them in the same directory. For large bandwidths sampled this feature, ppm error correction, has an unnoticably small effect but I wanted to add it anyway.

To call the spectrogram/log generator by itself for 431.2 Mhz at 2.4MS/s with a gain of 30 and frequency correction of 58 PPM use it like,

```
python graphfreqs.py 431200000 2400000 30 58
```

I've disabled the matplotlib (python) per frequency spectrogram plots for frequencies over 1 Ghz because there's not much going on up there. Also, the x-axis ticks and labels become inaccurate for some reason.

### Logs and forma

The signal strength logs, named by frequency (e.g. 53200000.log), use unix time and are comma seperated with newlines after each entry. In order of columns it is: unix time, relative signal level, gain in dB, PPM correction

```
1345667680.28 , -34.65 , 29 , 57
1345667955.59 , -34.67 , 29 , 57
1345668004.37 , -34.55 , 29 , 57
1345668110.06 , -33.88 , 29 , 57
```

It also generates a log file with all frequencies for use with gnuplot, all.log. This file has unixtime first, then frequency, then gain and ppm error.

```
1347532002.5 5200000 -14.84 29.0 58
1347532004.88 53200000 -17.84 29.0 58
1347532007.04 5400000 -17.98 29.0 58
1347532009.04 55600000 -19.78 29.0 58
1347532009.04 55600000 -24.04 29.0 58
1347532011.02 56800000 -24.04 29.0 58
1347532012.98 58000000 -26.21 29.0 58
1347532012.98 58000000 -25.12 29.0 58
```

### radioscan.p

The radioscan.pl script is used to automate calling graphfreqs in arbitrary steps. To generate plots and signal strength for 52 Mhz to 1108 Mhz with a gain of 30, sample rate of 2.4MS/s, and an interval between center frequencies of 1.2 Mhz, call it like,

\$ perl radioscan.pl -flist "52-1108,1248-2200" -g 30 -r 2400000 -s 1.2

### cli switches/options

```
-flist "52-1108,1248-2200" :: sets of frequency ranges to scan.
-g 30 :: gain
-c 2400000 :: sample rate
-d1 /path/here :: path to where the scripts are if now pwd
-d2 /path/here :: path to the directory to put logs, plots, gallery
-c 751 :: LTC Cell scanner frequency of fset correction, takes freq in Mhr of base cell
-c 751 :: LTC Cell scanner frequency of fset correction, takes freq in Mhr of base cell
-c 751 :: LTC Cell scanner frequency and the series charts for every freq
-c 751 :: turn on quaplot time series charts for every freq
-c 752 :: series first full range spectral map using all.log
-mr "52-1108,1248-1400,1900-2200" :: set of frequency ranges to plot as a another spectral map
```

```
Because I can use the default directories I keep it running like the below, but anyone else should make sure to set -d2.

$ while true; do perl radioscan.pl -flist "52-1108,1248-2200" -g 30 -c 2400000 -s 1.2 -w -c 751 -p -m -mr "52-1108"; sleep 1; done;

Running pyrtl graphfreq batch job 52 to 1108 Mhr at 1.2 Mhr spacings.

Using ITE Cell Scanner to find frequency offset from 751 Mhr station...Found Elonics E4000 tuner

Generating pyrtl graphfreq because the state of the station of
```

# Tuner/USB freeze solution with unplugging

starting rsync...

edit: as of Jan 5th 2013, librtlsdr has added soft reset functionality

Since graphfreqs.py's initializing and calling of rtl-sdr happens so frequently there are sometimes freezes. To fix these the USB device has to be reset. In the past I would accomplish this by un and re-plugging the cord manually. But that meant lots of downtime when I was away or sleeping. So, I've added in a small C program to the peri script using Inline::C that exposes a function, resetusb(). It is used if the eval loop around the graphfreqs call takes more than 10 seconds. This means you need Inline::C to run this script. To look at the original C version with a good explanation of how to use it click here.

```
good explanation of now to use a security of the security
```

My rtlsdr receiver + w/gnuradio implementation of the 11 GHz VSRT solar interferometer

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As far as I understand it, the VSRT design is a subset of intensity interferometer that uses the frequency error between multiple 11 GHz satellite TV "low noise downconverter block" (LNBF) clocks to create a beat frequency in the total power integrated. I am basically copying the MIT Haystack Very Small Radio Telescope (VSRT) but replacing the discrete component integrator and USB video input device with an rtlsdr dongle. The idea is to spend as little on hardware as possible.

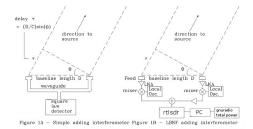
With modern LNBF the error between same model parts is about 30 ppm which results in beat frequencies of ~100 KHz at the 10 GHz of the mixers. With this kind of front-end there are no nulls but the fringe modulation can still be read out as variations in count of histogram bins that contain the beat frequency (in the total power fit). This intensity measurement proxy traces out the the envelope of the fringes and varies as a sinc function of distance between antenna. Knowing this and the distance can give you high angular diameter and position measurements of very bright radio sources.

```
To very origin tand sounces.

75 % z 18° satellite dishes w/mourts shipped
810 % KU LNFB (Fill21 8-2, -30pm error, RDA3560 w/27Mhr xtal.)
810 tisds receiver (F2800 reck, -30pm error)
810 power combiner (cheaper ones work too)
810 power combiner (cheaper ones work too)
820 coaxial power supply (LPF 12808) + diodes
820 loofer.Ref quadshield + F connectors
8130 Two Dish Fosition Motors (HH90)
800 DVB-S RI card (Skystar 2, DiShqC 1.2)
810 LBRG 2.1.8 mitch card (Skystar 2, DiShqC 1.2)
810 ELRG 2.1.8 mitch card (Skystar 2, DiShqC 1.2)
810 ELRG 2.1.8 mitch card (Skystar 2, DiShqC 1.2)
810 ELRG 2.1.8 mitch card (Skystar 2, DiShqC 1.2)
810 ELRG 2.1.8 mitch card (Skystar 2, DiShqC 1.2)
810 ELRG 2.1.8 mitch card (Skystar 2, DiShqC 1.2)
```

### Historical and other context.

For a detailed mathematical explanation of VSRT see MIT Haystack's VSRT Introduction. There is also a thread on the Society for Amateur Radio Astronomers list discussing the VSRT design. The more general concept of intensity interferometry, where you correlate total power instead of frequency, was originally developed by Hanbury-Brown & Twiss. Roger Jennison was around too. "The Entry Years of Radio Astronomy," Reflections Fifty Years after Janskys Discovery" by WT Sullivan (2005) is an excellent source about Hanbury-Brown and Twiss's side of it. The chapter "The Invention and Early Development of The Intensity Long Basseline Interferometry," (pdf). It covers not only the technical concepts but also historical contents, detailed hands-on implementations, and other personal anectdotes. And check out Jennison's book "Radio Astronomy" (1966)) as he invented the process of phase closure which uses a third antenna signal combined mathematically to recover some of the missing phase information. Arranged in a triangle of projected baselines the phase errors cause equal but opposite phase shifts in ajoining baselines, canceling out in the "closure phase". The MIT Haystack groups managed to resolve individual sunspots groups moving across the solar disk using with the technique with the VSRTs.



"An interferometer is an instrument that combines two signals (normally from two detectors) in a manner that the signals interfere to produce a resultant signal. The resultant signal is usually the vector sum of the two signals, but in some cases it is the product or some other mix. The traditional interferometer, usually studied and analyzed in physics courses, combines the two signals in a way that both amplitude and phase information are used. By varying the positions of the two detectors, it is possible to synthesize an effective aperture that is equivalent to the separation of the detectors and to reconstruct the impinging wavefront, thus providing significant information about the extent and structure of the signal source. The traditional phase-sensitive interferometer requires retention of the signal phase at each detector — the phase-sensitive interferometry technique will not be discussed in detail here."

"A special case of the interferometer is the intensity interferometer, which performs an intensity correlation of signals from the two detectors. Although in the intensity interferometer the phase information from the two antennas is discarded, the correlation of the two signals remains useful. Aperture synthesis is not practical, but some important source characteristics may be determined."

I think the VSRT is a special case of intensity interferometer where you don't try to align samples by time after recording. Instead you just look for the baseline distance sine pattern in total power at the beat frequency of the unsynchronized clocks.

# Implementation so far.

So far I've only done it with manual pointing screwed to a board. The interferometry correlation is done with a satellite tv market stripline power combiner at the intermediate frequency (IF, ~950-1950 MHz) and then an rtlsdr dongle is used to measure the total power of a 2.4 MHz bandwidth of the intermediate frequency range. I use a gnuradio-companion flowgraph to take the total power and then do a fourier transform of the total power. In this fourier transform the fringes show up as a modulation of the count in the FFT bins which correspond to the difference in frequency between the two downconverters. In my case this is about ~100 KHz.

In the Haystack VSRT memos a line drop amplifier, or two, are sometimes put behind the respective LNBF IF coax outputs or the power combiner. With the rtlsdr dongle and relative short (<10m) baselines of RG6 this isn't required.

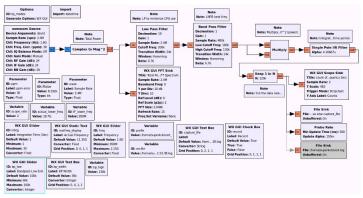
The GUI allows for setting the exact 2.4 MHz bandwidth of the IF range to sample and the total power FFT bin bandpass to where and what the LNBF beat frequency is. The file name is autogenerated to the format,

prefix + datetime.now().strftime("%Y.%m.%d.%H.%M.%S") + ".log"

The time embedded in the filename is later used by a perl script, vsrt\_log\_timeplot.pl, which converts and metadata tags the binary records to gnuplot useable text csv format for making PNG plots.

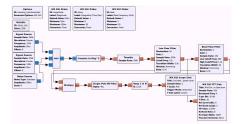
# Download

total power modes (tp-modes.grc)
 vsrt\_log\_timeplot.pl



Who else helped

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I consulted with patchyonbraun a lot for the software/gnuradio side. He gave me an example of how to use the WX GUI Stripchart and I would not have guessed I needed to square the values from the beat frequency bins after the first squaring for taking total power. He made a generic simulator for dual free running clocks LNBF intensity interferometers. You don't even need to have an rtlsdr device to run it; only an up to date install of gnuradio. It is an easy way to understand how to do interferometry without a distributed clock signal.

patchvonbraun's: simulated-intensity-interferometer.grc

# Physical



With this setup on a 1 meter baseline and a intermediate tuning frequency of 1.6 GHz IF (10700 MHz+1600 MHz-950 MHz)= 11350 MHz) the main beamwidth would be about 70% (c/11GHz)/1m), or 1.9 degrees. This does not resolve the solar disk (~0.5 deg) during drift scans. I have been told that the magnitude goes down in a SINC pattern as you widen the baseline and approach resolving the source but I will not resolve the sun initially. In the VSRT Memos "Development of a solar imaging array of Very Small Radio Telescopes" a computationally complex way to resolve individual action regions is done with a 3rd dish providing "phase closure" in the array on a slanted north-south baseline in addition to the existing east-west baseline. I try to point my dishes so that the Earth is passing the sun through the beam at ~12:09pm (noon) each day. To aid in pointing a cross of reflective aluminum tape is applied center of the dish. This creates a cross of light on the LNBF feed when it is in the dish focal plane and the dish is pointed at the sun. The picture below is from later in the day, the one of the left shows the sun drifting out of the beam as it sets. I made my LNBF holders out of small pieces of wood compression fit in the dish arm. There are grooves for the RG6 coax to fit ground out with a rotary tool. The PVC collars have slots cut in the back with screws going into the wood to set the angle.



The screenshot shows a short run near sunset on an otherwise cloudy day. The discontinuities are me running outside and manually re-pointing the dishes. But it does highlight how the beat frequency of the 2 LNBF varies as they warm up when turned on. It starts down at ~90 KHz but within 10 minutes it rises to ~115 KHz. After it reaches equilibrium the variation is ~+1 KHz. I could change the existing 80-120 KHz bandpass to a 110-120 KHz bandpass and have better sensitivity. But that bandwidth is something that has to be found empirically with each LNBF pair and set manually within the GUI for now.

patchvonbraun said it was feasible to identify the frequency bins with the most counts and that there was an example within the simpla\_ra code,

"You could even have a little helper function, based on a vector probe, that finds your bin range, and tunes the filter appropriately."

The below close up of indoor testing showing how everything is connected on the rtlsdr side showing the power injector, e4k based rtlsdr (wrapped in aluminum tape), and the stripline based satellite power combiner for correlation. The two rg6 quadshield coaxial lines going from the power combiner to the ku band LNBF are as close to the same length as I could trim them. I use a I amp 18v power supply and coaxial power injector to supply power to the LNB and any amplifiers. This voltage controls linear polarization (horiztonal/vertical) and it can be changed by putting a few I amp 1N4007 in series with the power line to drop the voltage.



Accessory scripts

tp-modes, grc produces binary logs that are pretty simple. The count of the LNBF beat frequency bins in the bandpass are saved as floats represented as 4 pairs of hexadecimal. When the integration time is set to the default 1 second then one 4 byte data point is written to the log every 0.5 seconds. I highly recommend not changing this for now. There is no metadata or padding. Here's a screenshot of a run using the utility "bless",

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```
File Edit View Search Tools Help
 咯 🖹 👲 | 🖫 🐑 🖒 | 💥 🖳 🖺 🖊 🎮
 vsrt 2013.06.14.12.03.00.log ×
                                                     Hexadecimal: 6A EB 87 35

Decimal: 106 235 135 (
                              Signed 32 bit: 898100074
                                                                    Decimal: 106 235 135 053
  Unsigned 8 bit: 106
                            Unsigned 32 bit: 898100074
   Signed 16 bit: -5270
                         Float 32 bit: 1.01268E-06
Float 64 bit: 8.22582026993166E-51
                                                                     Octal: 152 353 207 065

Binary: 01101010 11101011 1000011:
 Unsigned 16 bit: 60266
                             ☐ Show unsigned as hexadecimal

✓ Show little endian decoding

                                                                     ASCII Text: j??5
                                     Offset: 0x290c / 0x7d97
                                                                    Selection: 0x2908 to 0x290b (0x4 bytes) INS
In order to convert the binary logs of 4 byte records into something gnuplot can parse I use a simple perl script,
```

```
my Sdata = '/home/superkuh/vsrt_2013.06.13.12.26.47.log';
my Sbytelength = 4; # floats (little endian)
my Somm records;
Sdata = $AKGV[U];
} else {
    print "you need to pass the log file path as an argument.";
    exit;
open(LOG,"$data") or die "Can't open log.\n$!"; binmode(LOG);
my $i = 0;
until (eof(LOG)) {
    my $record;
    my $decimal;
    read(LOG, $record, $bytelength) == $bytelength
        or die "short read\n";
    3decimal = unpack($format, $record);
    printf("$i,\tSdecimal\n", $decimal);
    3i+;
```

Now I have the filename which gives the time the gnuradio-companion grc file started running. This is not the time I hit the record button and started logging. The offset is a second or two. I gnoring that, it is possible to use the start time encoded in the log file name to figure out when a narticular measurement was taken. To do that I have to know the interval between entries saved to the binary log.

```
((30124-29644)/4)/60 = 2
$ date 66 ls -1 /home/superkuh/vsrt_null.log 66 sleep 60 66 date 66 ls -1 /home/superkuh/vsrt_null.logFri Jun 14 13:44:36 CDT 2013
-rw------- 1 superkuh superkuh 8 2013-06-14 13:44 /home/superkuh/vsrt_null.log
-rw--r---- 1 superkuh superkuh 488 2013-06-14 13:45 /home/superkuh/vsrt_null.log
```

To know what time a log record corresponds to, take the time from the filename and then add 0.5 seconds \* the index of the 4 byte entry in the binary log. This should be possible to write into the until loop so it outputs time instead of just index \$i\$. The below example is a hacky version of my log parser that does just this. Here's an example output.

```
# The simplest possible gnuplot plot using this program's output.
# ./vsrt log timeplot.pl /home/superkuh/vsrt 2013.06.14.12.03.00.log > whee2.log
# gnuplot> plot "./whee2.log" using 1:2 title "VSRT Test" with lines
 my Sdata = '/home/superkuh/vsrt_2013.06.13.12.26.47.log';
my Sbytelength = 4;
my sformat = """; # oops, not this unsigned 32 bit (little endian)
my Sformat = "f"; # float my Smur_ecoords;
 if ($ARGV[0]) {
    $data = $ARGV[0];
Sdata = SARSV[U];
} else {
    print "you need to pass the log file path as an argument.";
    exit:
 my $dt; # declare datetime variable globally extracttime($data); # $dt now has date object.
open(LOG, "$data") or die "Can't open log.\n$!";
binmode(LOG);
umy 5i = 0;

until (ecf(LOG)) (

until y Stecord;

my Stecord;

my Sdecimal;

read(LOG, Stecord, Sbytelength) == Sbytelength

or die "short read\n";
               # This is a stupid/fragile way to deal with daterime
# not having enough precision. It only works if the
# record to record interval is always 0.5 seconds.
my Srecordisine = Sdt->epoch();
if (0 == Si % 2) {
    printf("Srecordisme.0.\tSdecimal\n", Sdecimal);
               } else {
    printf("$recordtime.5,\t$decimal\n", $decimal);
               $dt->add( nanoseconds => 500000000 );
$i++;
```

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```
my (Spear, Santh, Sday) = split(/\./, Syear month_day);
Stimes // //// spear month_day);
Sdt = DateTime->new(
year => Syear,
month => Syear,
month => Syear,
hour => Shour,
hour => Shour,
accod => Saccod,
stime zone => 'America/Chicago',
;
Sdt => Set time zone ('UTC');
return [;
```

Now I just have to make up a good gnuplot format and integrate the calls into the perl script.

### Computer controlled pointing, mechanical and software differences

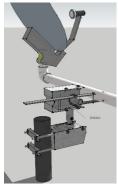
MIT Haystack's memo #9 photo of their dish pointing system



Figure 2. Azimuth/elevation mount using 2 DiSEqC SG2100 motors

Manually repositioning the dishes swamps out the signal of interest as the target leaves the beamwidth. For any decent measurements I need computer controlled pointing. This means the Haystack idea of two coupled Diseqc 1.2 compatible motor positioners mounted one on the other. In their design both dishes are mounted on a single PVC tube hooked to one of the positioners with a metal extension. My satellite dish mounts can't rotate like theirs so I'll have to modify this design a bit. They use a serial relay to "push" the buttons on a physical Diseqc 1.2 motor controller remote. That seemed a bit convoluted to me. I bought a SyStart DVB-Sp. eig card and undered relinux send raw by calling staffne which accesses the limix DVB interface. It has both a GUI and cli interface. Unfortunately xdipo cannot send through Diseqc switches. I had to add manual motor commands to tunes-s2 which did support switches but not manual motor commands. This version which supports manual stepping mode is available at <a href="https://github.com/superkul/tune-s2-stepping">https://github.com/superkul/tune-s2-stepping</a>.

Another alternative Diseque motor controller I didn't persue would be using a 192 KHz USB soundcard and the <u>DiSEqC Audio Generator</u> software from Juras-Projects. The documentation for the hardware side of the audio generator is 404 now, but <u>Juras responded to an email of mine with the schematics attached.</u>



Since the bent motor shafts that came will my motors looked really difficult to drill through I'd use straight hex holed shafts to make everything mechanically simpler. I found <a href="http://www.reidsupply.com/sku/HHS-18/">http://www.reidsupply.com/sku/HHS-18/</a> and ordered a couple. Unfortunately my measurements of the dish motor shaft flat-to-flate size were off. The Reid hex holed shaft hole is just a tiny bit too large. This was easily fixed by wrapping a couple turns of masking tape around the shaft to increase the diameter. This is often how fishing rod handles are made. I also encountered this construction technique on Jarrod Kinsey's CQJ laser pages.

The hardest part of all this is drilling an 8mm hole precisely normal to the curved outside surface of the hex hole shaft. The first step is to flatten the area with a hand file. This took me about 10 minutes. I had previously ordered and received two carbide drill bits, one small to sub-drill the initial hole and then one 8mm for the final hole. A drill press and small vice are quired to actually drill the holes. And even then it's really tricky. My first two attempts resulted in holes not quite normal to the surface of the hex flat. I could only use roll or taper pins to secure the shaft. Luckly I bought 2x shafts just in case.

I also had to drill 4 additional 8mm holes in the 2x satellite dish motor mounts to make holes for level mounting instead of at a till. The VRST guys got lucky with their sat motor mounts having a long slot.

The diameter of easily available PVC is slightly to small for the dish mounting clamp. This is remedied like the motor shafts; by wrapping wide masking tape to size and optionally epoxy coating/sanding it.



The dish motors used in the VSRT project were Stab HH90. These have come down in cost since the VSRT memos were written and are still widely available.

```
065 Stab HH90 dish pointing motor #1
065 Stab HH90 dish pointing motor #2
```

In order to control these motors a system to send <u>DISEGC 1.2 commands</u> is needed. The first option would be to faithfully replicate the VSRT implementation. They do it in a rather roundabout way but at least it is tested and known to work with their software. Unfortunately the specific hardware used has become rare, is mostly shipped from overseas, or is expensive.

```
045 STAB MPO1 Positioner Control #1
045 STAB MPO1 Positioner Control #2
0573B MPO1 Positioner Control #2
058T Java software (windows)
000 VSRT Java software (windows)
000 VSRT Java software (windows)
000 Two "5" *9" *0.125" thick steel plates
020 Two 15" *1" *0.125" thick aluminum counterweight arms
020 Shipping for the metal parts.
```

My chosen method of HH90 motor control is a single DVB-S card under linux with DVB API 5.x w/ my modified tune-s2 and optionally xdipo. This can be combined with a DiSEqC switch to scale to control of multiple motors relatively cheaply. I do sun alt-az position calculation by using a small pysolar python script. I have not yet completed the scripts to turn alt-az positions of the sun at my location into motor step commands. Hopefully I can use some of the USAL fuctions in tune-s2 for that.

```
066 Skystar 2 HD pci card
010 DisEqC switch
0 tune=22 and xdipo DVB control software (linux)
020 . 753" Hex-Holed Sleeve
10 5" wide masking tape
```

Both require PVC pipe, tools like drills, 8mm drill bits and smaller sub-drill bit, hand saws, files, and potentially a welder (though liberal J-B Weld would probably work).

# Diseqc switches problems and solutions.

It turns out that xdipo alone cannot deal with motors behind Diseqe switches. This means it can only control one Diseqe motor at once. Controlling two would require 2x Skystar 2 pci cards. Luckily there are other options. CrazyCar's <a href="mailto:supports">mmess2</a> supports Diseqe switches and addressing It normally only provides for motor commands using the <a href="mailto:support supports">Skystem visit in support</a> supports Diseqe switches and addressing It normally only provides for motor commands using the <a href="mailto:support support sup

This gutted version of tune-s2 for manual motor commands is available at:

# https://github.com/superkuh/tune-s2-stepping

The functions I added are basically just look up arrays with Diseqc bus commands for different steps in the clockwise or counter-clockwise directions. In Diseqc the packets have 4 sections. Check out the Diseqc Bus Functional Specification (pdf) for a better explanation with more detail

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```
FRAMING P ADDRESS P COMMAND P DATA P
```

The first, "Framing" byte represent if the command is from the receiver or diseqc device and wether it needs a reply. For my table these are all just "EO" which means it's a packet from the receiver with no response required. Most commands are EO but it goes up to E7.

The second, "Address" specifies which types of Diseqc devices should listen (ex: LNB, switch, motor, polarizer). For motors this is "32"

The third is "Command". This is a huge list of values of which only "68" and "69" are relevant. They are "Drive Motor East" and "Drive Motor West" respectively. The "Command" byte is only relevant to their specific devices specified via the "Address" byte.

The remaining bytes of the packet are "Data" and how they're interpreted depends on the "Command" bytes specifying a specific type of command. For motor movement there are three options. "00" makes the motor turn until a Diseqe stop command is sent. The second mode is positive values for the bytes, "01" to "7F". They represent an amount of time to turn the motor. Or by specifying negative byte values "80" to "FF" the motor is rotated a number of steps. This last is best and detailed in the <u>Positioner Application Note</u> (pdf) with an except below,

The number of steps to make is given by the additional count needed to make the parameter byte reach zero (or overflow to zero if the byte is considered as unsigned). Thus the byte 'FF' (hexadecimal) requests only one step, FE' two steps, and for example F9' requests 7 steps

With my motors each step corresponds to about 0.1 deg. Using this information I made up a table of Diseqc packets for each rotation direction.

For addressing specific ports of the Diseqc switch tune-s2's normal functions are used. They are called before the motor position commands are sent. Usage of the modified tune-s2 is pretty simple. The only differences are two new cli switches and not needing to give it tuning parameters.

-step-east -step-west

### They each take any value from 0 to 10 like.

./tune-s2 -step-west 0 -committed 1

This would cause the satellite dish motor on port 1 of the Diseqe switch to step 1 position counter-clockwise. To send the same command of stepping 1 position counter-clockwise to the other motor

/tune-s2 -step-west 0 -committed 2

The stepping argument values 0 through 10 are mapped on a fairly arbitrary set of actual steps. This results from just doing array index look ups in the above packet tables.

```
0->1
1->2
2->3
3->4
4->5
5->10
6->20
7->30
8->40
9->50
10->100
```

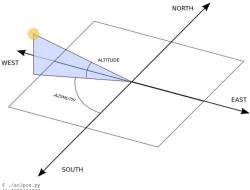
NAME OF THE PARTY OF THE PARTY

### Calculating solar position and using that to decide how many steps to step per axis

Figuring out where the sun is in the sky in terms of an alt-az format is made simple by pysolar. Figuring out how to turn that position into sequences of steps on the motors is much, much harden

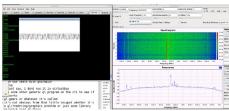
```
## /mar/him/env sython import pymoin import pymoin ar import datetime d = datetime.datetime.utcnow() lat = 40.0 long = 90 long
```

These values are relative to the pysolar reference frame which is given by their diagram,



My setup is pointed directly south. So for this example time that means I need to calculate the number of steps required to turn the (top) altitude motor 41.5 up from level and the (bottom) azimuth motor 35.8 degrees to the left (east).

# Decoding Pager Data with multimon and/or gnu radio receivers



The hardest part of this is figuring out what kind of pager system you have. I spent a long time trying to decode the local FLEX pager system with decoders that did not support it.

Written by Thomas Sailer, HB9JNX/AE4WA, multimon (multimon tar.bz2) supports decoding a large number of pager modulations. FLEX is not one of them. Scroll down for FLEX

On June 29th 2012 dekar told me about his updated fork of multimon, multimonNG, with better error correction and more modulations supported. As of right this instant those on 64bit linux should just use the existing makefile and \*not\* qmake or qt-creator to compile it. For the windows users (or anyone wanting more info) there's a precompiled version and blog post. Make sure to disable all the demodulators you don't need. I think especially ZVEI is quite spammy. This and this is what poesag sounds like if you're wondering.

When I originally started playing and wrote this there were only a couple options for rtlsdr receivers to use with the multimon decoder. I used patchvonbraun's multimode to save .wavs and dekar's pager example GRC I modified for OsmoSDR sources linked below for raw, real time decoding. Lately (as of late 2012/13) a large number of receivers have been released that don't depend on GNU Radio. rtl\_fm is one and there's an example usage below.

# real time decoding rtl\_fm

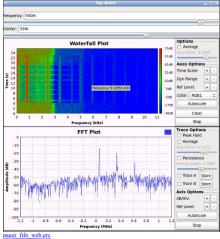
```
rtl_fm -f 930.353e6 -g 100 -s 22050 -1 310 - |multimon -t raw -a POCSAG512 -a POCSAG1200 -a POCSAG2400 -f alpha /dev/stdir
```

# real time decoding w/dekar's pager\_fifo

Dekar's multimonNG, a fork with improved error correction, more supported modes, and \*nix/osx/windows support. In the screenshots below the signal is not pocsag. I thought it might be zwei but now I'm not so sure it's even pager data. Test samples of pocsag that Dekar links on his

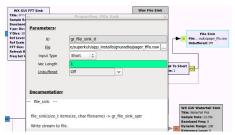
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# blog decode just fine

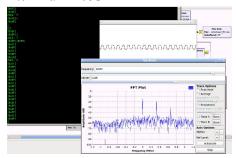


mkfifo /tmp/pager\_fifo.raw ./multimonNG -t raw /tmp/pager\_fifo.raw gnuradio-companion pager fifo web.grc

In order to decode the pager data in real time you should use a first-in first-out file (fife). Dekar's pager\_fifo is designed to do that but you'll need to set the correct file paths for the File Sink yourself. In the copy downloadable here the File Sink's path is set to "/tmp/pager\_fifo.raw". You should be able to run it without editing once you've made that fifo. Make sure to start multimon reading the fifo before you begin GRC and execute the receiver.



In my personal copy of dekar's pager fifo the file and audio sinks are enabled while the waterfall, wav, and other sinks are disabled. To enable the disabled (grey) block select them and press 'e' ('d to disable). The audio sink is set to pulseaudio ("pulse").



# FLEX Pagers

Unfortunately it turned out my local pagers were all using FLEX, and so not supported by any of the above software. But the procedures might still be useful for someone. Decoding FLEX can be done with the software PDW, but it is windows only. In GNU Radio there is additionally grager, which is supposed to support flex, but many implementation scripts for it are GNU Radio 3.6.5 or older and getting stuff to work with 3.7 requires namespace changes, motherar's flex\_hacker' is one of these. Since the relsdr receivers can but shouldn't do 3.125 MS/s, like flex\_hacker' for flut\_flux, what they use natively for the bandwidth, and so decimation, and pretty much everything else have to be re-written too. I've attempted to start this and you can see a copy\_here.

A couple days after I wrote the above paragraph zarya came on ##rtlsdr on freenode and mentioned his rilsdr supprting FLEX decoder written months before. It is easy to use and works great! This script runs at a 250 KS/s sample rate and decodes 12.5 KHz channel only. Internally it uses gnuradio's optfir to generate low pass taps that wide to use with a frequency xlating FIR filter. It then passes that to gr-pager's flex\_demod.

later: argilo (Clayton Smith) has also put together an osmosdr source based gr.pager flex decoder for his GNU Radio tutorial series.

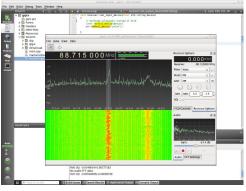
The below output is heavily censored and edited to avoid disclosing or reproducing sensitive information but it gives you an idea of the type of messages.

```
qit clone https://qithub.com/zarya/sdr
cd sdr/receivers/flex/
./tl_flex.ox.py = f 929.56M -=rx-qain=37.2
linux; GNU C++ version 4.6.3; Boost_104800; UHD_003.005.004-149-gc357a16e
IRRA NW 400 THE AND TH
```

I tried for over a year before successfully decoding local pager signals. Now that I have I think it is a bad idea. There is far much too much private information in cleartext. I don't plan to try again

(old) gqrx install notes

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When I wrote this up the original version by csete didn't support the hardware yet but mathis\_, phirsch, Hoernchen, and perhaps others I've missed from ##rtlsdr on freenode had added librtlsdr support to gqrx; their repos are still listed by commented out. These days esete has added in rtlsdr support so you can use his original repository.

```
git clone https://github.com/csete/gqrx.git
od ggrx

† on Ubuntu, sudo apt-get install qtcreator , if you don't have it.
qtcreator ggrx.pro  

† press the build button (the hammer)

† Avoid qtcreator doing it manually.
qmake
```

### Use with Ubuntu 10.04 and distros with old Qt < 4.7

You will almost certainly not get this error. But, someone might, so I'm leaving it here to be indexed.

If you're like me and run an older distribution then your Qt libraries will be out of date and lack a function required for generating the name of the files to be saved when recording

Initially I thought it was a quereator thing so I tried to get more information by doing it manually,

make get+-c-pipe-O2-l/usr/local/include/gnuradio-l/usr/local/include-l/usr/local/include-l/usr/local/include-l/usr/local/include/gnuradio-D\_REENTRANT-D\_REENTRANT-l/usr/include/glubusb-1.0 -Wall -W-D\_REENTRANT-DQT\_NO\_DEBUG-DQT\_

To get it to compile on these systems you'll have to do the below. (edit: This little change is now added into phirsch's.)

Ubuntu 10.04 has old Qt libs and gqrx uses a function call not in them. So, while I was waiting for Qt 4.74 to compile I decided to try a hack. I removed that function call with a static string of text. [edit] I later found comparable functions for Qt 4.6 and older.

If you are using accreator like the does suggest you can double click on the error and go to the line. If not, it was in /Sources/qtgui/dockaudio.cpp replace,

```
void DockAudio::on_audioRecButton_clicked(bool checked)
{
   if (checked) {
    //FINGE: option to use local time
    lastAudio = QDateFine::currentDateFineUtc().toString("gqrx-yyyy%%dd-hhmmss.'wav'");
```

### With something like this

```
void DockAudio::on audioRecButton clicked(bool checked)
    if (checked) {
    // FIXME: option to use local time
    // use functions compatible with older versions of Qt.
    lastAudio = QDateTime::currentDateTime().toUTC().toString("ggrx-yyyy0Mdd-hhmmss."wav"");
```

And it'll compile and run correctly on my specific machine.

# Compiling LTE Cell Scanner and LTE Tracker on Ubuntu 10.04

Before starting make sure to have a fortran compiler, FFTW, BLAS, and LAPACK libraries installed from the repositories.

do apt-get install automake autoconf libtool libfftw3-3 libfftw3-dev gfortran libblas3gf libblas-dev liblapack3gf liblapack-dev libatlas-base-dev

If you're using 12.04 just follow the instructions on the github page and everything is trivial. For 10.04 (lucid) users the the initial hurdle is emake. LTE Cell Scanner requires cmake 2.88 and Ubuntu 10.04 only has 2.8 the finding of BLAS and LAPACK libraries will fail like

```
CMake Error at CMakeLists.txt:1 (CMAKE_MINIMUM_REQUIRED):

CMake 2.8.4 or higher is required. You are running version 2.8.0.
```

Until you open CMakeList.txt and change the version number on first line to 2.8.0. After fixing that the BLAS and LAPACK issues come in,

```
cmake .

- Found ITPF: /usr/lib64/libitpp.so
CMake Error at /usr/share/cmake-2.8/Modules/FindBLAS.cmake:45 (message):
FindBLAS is Fortran-only so Fortran must be enabled.
Call Stack (most recent call first)
CMakeLists.txt:29 (FIND_BCKRGE)
```

You can see my installation notes before I figured it out. To fix it I searched for people complaining of similar problems on other projects and then replaced my \*system\* files with theirs, FindBLAS.cmake

```
sudo cp /usr/share/cmake-2.8/Modules/FindBLAS.cmake /usr/share/cmake-2.8/Modules/FindBLAS.cmake.bak sudo cp FindBLAS.cmake /usr/share/cmake-2.8/Modules/
```

LAPACK will also fail this way. I used this arbitray cmake file, <a href="http://code.google.com/p/qmcpack/source/browse/trunk/CMake/FindLapack.cmake?r=5383">http://code.google.com/p/qmcpack/source/browse/trunk/CMake/FindLapack.cmake?r=5383</a>. And <a href="http://code.google.com/p/qmcpack/source/browse/trunk/CMake/FindLapack.cmake?">http://code.google.com/p/qmcpack/source/browse/trunk/CMake/FindLapack.cmake?r=5383</a>. And <a href="http://code.google.com/p/qmcpack/source/browse/trunk/CMake/FindLapack.cmake?">http://code.google.com/p/qmcpack/source/browse/trunk/CMake/FindLapack.cmake?r=5383</a>. And <a href="http://code.google.com/p/qmcpack/source/browse/trunk/CMake/FindLapack.cmake/trunk/source/browse/trunk/source/browse/trunk/source/browse/trunk/source/browse/trunk/source/browse/trunk/source/browse/trunk/source/browse/trunk/source/browse/trunk/source/browse/trunk/source/browse/trunk/source/browse/trunk/source/browse/tru

sudo cp /usr/share/cmake-2.8/Modules/FindLAPACK.cmake /usr/share/cmake-2.8/Modules/FindLAPACK.cmake.bak sudo FindLAPACK.cmake /usr/share/cmake-2.8/Modules/FindLAPACK.cmake

After fixing the cmake issues compile and install the latest 11++ (ITPP 4.2). Make sure to completely remove the old ITPP 4.0.7 libraries from the Ubuntu repository. When LTE Cell scanner compiles you can go back and restore the .bak cmake files. The rate of scan is about 0.1 Mhz per 10 seconds.

```
DET 10 SECURIS.

(*CellSearch -v -s 75le6 -e 75le6
LTE CellSearch v0.1.0 (release) beginning
Search frequency: 751 MHz
FPM: 100
correction: 1
Found Elonics E4000 tuner
Maiting for AGC to converge...
Examining center frequency 751 MHz
Calculating PSS correlations
Searching for and examining correlation peaks...
**Perceted a cell 1
                                                 Searching for and examining correlation
Detected a cell!
cell ID: 414
RX power level: -17.0733 dB
residual frequency offset: 43592.8 Hz
     AX power level: -1.0/35 das las residual frequency offset: 43592.8 Hz level: -1.0/35 das level: -1.0/35 das
```

Both positive and negative frequency offsets happen, but rarely in the same dongle.

LTE Tracker I haven't used as much yet (recently released) but it is included in the github repository cloned initially and should be compiled as well if you did the above. Check out the authors site for videos of it's use since an ascii paste of the neurses like interface wouldn't tell you much. But... the start looks like this,

```
/LTR-Tracker - 7 751e6
LTE Tracker - 7 751e6
LTE Tracker v1.0.0 (celease) beginning
Search frequency: 751 MHz
FPM: 120
correction: 1
Found Rafael Micro R200T tuner
Calibrating local oscillator.
Calibrating local oscillator.
Residual frequency offset: -48937.5 Hz
Residual frequency offset: -48937.5 Hz
```

Стр. 16 из 21 20.11.2020, 23:11 New correction factor: 0.99993484110674779597 Searcher process has been launched.

### Slightly altered GNU Radio Companion flowcharts

"The GUI stuff in Gnu Radio was rather an afterthought. Nobody really expected that you'd use it to build actual applications, but rather just use it as a way of making "test jigs" for your signal flows."

This section is my notes on how I made basic examples work, and how I edited those examples in very simple and often broken ways. Also, since gqrx, multimode, and other intergrated receivers came out I don't see any need to update these as things change. Most of this is very old.

While there are links to the originals in the summaries, these descriptions are of the versions modified by me; usually just sample rate and GUI stuff. While the sample rate or tuner width I set may be some large number; it'll become obvious what the limits of each other as you scan about and see the signal folding or mirroring. Using sample rates above 2.4 MS/s with rtlsdr is not recommended. It \*does\* create aliases all over. If you're using GNU Radio 3.7 don't even bother trying with any .grc files hosted here.

- FM:

- patchvonbraun's simple (stereo) fm receiver harder setup, best reception, best sound, 2.048 MS/s, +-600Khz fine tune
   imdfs fm receiver easy setup, good reception, good sound, 3.2 MS/s, +-600Khz fine tune
   superchuls offset fm receiver easy setup, okay reception, okay sound, 2.8 MS/s, +-900Khz tune, +-50Khz fine.
   2h20's beginner fm (mono) receiver easy to understand, easy setup, okay sound, 2.8 MS/s, no fine tune
- SSB:
  - OZ9AEC's SSB Receiver SSB rx and record to disk, seperate playback script. 1 MS/s, +-1k fine tune.

If it comes with a python file, try that first before generating one from the GRC file. When tuning, make sure to hit enter again if it doesn't work the first time or tunes to the wrong frequency. Always hit autoscale to start, and for FFT displays try using the "average" settings. I have set all audio sinks to "pulse" (pulseaudio) instead of say, "hw:/0,0" (ALSA). You might have to change that. To get a list of hardwareuse "aplay +". That"[ll show the various cards and devices. Use the format, "hw:X,Y" where "hw:CARD=X,DEV=Y". Some flowwhats have variables for it, others put it directly in the Audio Sink element, I fly ou hear something interesting you can try companies from <a href="http://www.kb/bukd.com/digital/">http://www.kb/bukd.com/digital/</a> or <a href="http://www.radioreference.com/">http://www.radioreference.com/</a> or <a href="http://

There are two ways to specify the use of multiple dongles. The first, correct, way is to set the "Num Channels" in the OsmoSDR Source block to "2" and then specify the device IDs in "Device Arguments" like, "rtl=0 rtl=1". Each specified device is seperate with a space from the

The not so correct but still working way is to use multiple OsmoSDR Source blocks with "Num Channels" set to "1" and each with it's respective "Device Arguments" field set to "rtl=0" or "rtl=1", or so on.

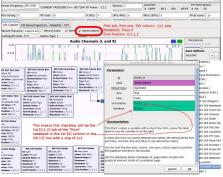
The OsmoSDR Source block has extensive help files at the bottom of it's properties if you scroll down.

To enable a block, select it and press 'e'. To disable a block select it and press 'd'. When disabled blocks will appear darker gray,

If you open a .grc file and it looks like there are blocks missing (red error highlights and no connections between them) then it is likely the name of the block changed during some GNU Radio update. If your install is more than a month or two old this often happens. Update GNU Radio

It's easier to type in 1e6 than 1000000 so use scientific notation when you can in variable fields. If you double click on an element in a flowchart it usually includes a helpful "Documentation." of most the variables to be set at the bottom. The GUI element grid position is a set of two pairs of numbers: "y,x,a,b," where the first pair "y,x," is position (y row, x column) and "a,b" is the span of the box. If you enter a Grid Position and it overlaps with another element it'll turn red and report the error and where the origin is of the element it overlaps with.

Use the Grid Position (row, column, row span, column span) to position the graphical element in the window



For RTL2832 Source the minimum sample rate is ~800KS/s, it's gr-baz(?) and generally not updated. Use OsmoSDR source. It's under "OsmoSDR", not "Sources" on the right panel). It has a IMS/s minimum sample rate. It's not recommended to use sample rates above 2.4M.

In older versions of gr-osmosdr and rtl-sdr I think automagic gain control (AGC) was on all the time so you didn't have to set the gain explicitly in the source in GRC. New versions require that and also require setting the chan 0. freq to something,

The dongles seem to have noise at their 0Hz center frequency so the best performance is from selecting a band 100-200Khz offset from the center (depending on signal type). patchvonbraun's simple fin rev is a great example of that.

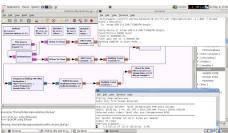
# patchvonbraun's simple\_fm\_rcv

(this summary is outdated) The best sounding software I've found for listening to FM is patchvonbraun's Simple FM (Stereo) Receiver. I don't think it is very simple; it includes many advanced FM specific features like extraction of the 19k (pilot) tone next to some commercial FM broadcasts. It used to do RDS, I hear, and older versions checked into CGRAN still have it, but it is removed for simplicitly in this version.

```
less rancommands and the second secon
```

# lindi's FM receiver

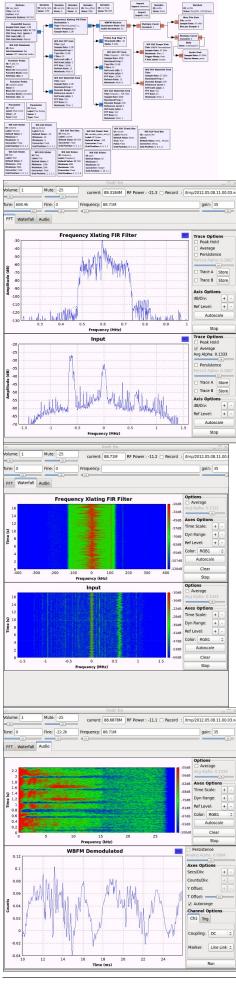
Original: http://lindi.iki.fi/lindi/gnuradio/rt12832-cfile-lindi-fm.grc, this was an example posted to ##rtlsdr by lindi. It used a file source which was decoded to way and saved to disk. Seen in the screenshot above



Modified: http://superkuh.com/rtl2832-cfile-lindi-fin\_edit.grg\_, had a frontend GUI and an increased sample rate. Right now the rate of the audio files saved out is... not very useful. But it sounds fine. Seen below

3.2 MS/s field of view, tune +-900Khz

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2h20's simple fm receiver

2.8 MS/s field of view, no fine tuning

h20 made available, with thorough explaination, a bare bones FM (mono) receiver to learn how to use GNU Radio. This was the first one I managed to get to work. Because the original h202's uses the RTL2832 Source and not the OsmoSDR Source you might experience tuner

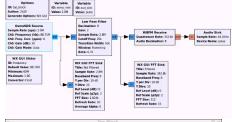
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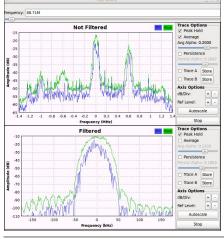
crashes if you scan too quickly. Make sure to un/replug in the dongle after these. It's best just to enter the frequency as a number.

[Be aware this section of this page was written many months ago when rtl-sdr was different and I had little idea of what I was doing. xzero has since manually added signal seeking to this example.]

My edit of 2h20's simple receiver does not add much, but I did replace the RTL2832 source with an OsmoSDR source to avoid tuner crashes. I also increased the sample rate to 2.8MS/s (to see more spectrum) and then increased the decimation in the filter from 4 to 8 to compensate so everything still decodes/sounds right. I also remove the superfluous throttle block.

• Modified 2h20's Mono FM Receiver (.grc)





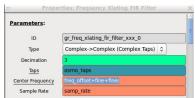
### my offset tuning + recording example

2.8 MS/s field of view, +-900Khz tuning, +-50Khz fine.

This takes parts from a bunch of the other example receivers and repurposes them in presumably incorrect but seemingly working ways. It is a basic example of how to offset the tuner 200khz away from the center to avoid the noise there. I started with 2h20's simple tuner's GUI framework and removed almost all of the content. I copied, with inaccurate trial and error changes of sample rate and filter offset, sections of the offset tuning and other advanced bits from simple\_fm\_rev and wim\_rx\_gec. The tuner is tuned +200khz. The freq\_xlating filter is tuned +200khz. The bandpass filter is specified in a variable,

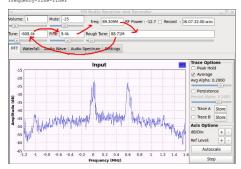
firdes.complex\_band\_pass(1.0,1.024e6,-95e3,95e3,45e3,firdes.WIN\_HAMMING,6.76)

The net result is that the frequency of interest comes out of the tuner 200Khz below DC, and the freq\_xlater "lifts it up" by 200Khz, and then it's bandpassed.



I also blindly copied the RF power display, a toggle for saving the audio files out to disk, and a+-900khz tuning slider from other receivers. I added a second 'fine' tune +-50Khz. This is done by setting the frequency of the Xlating FIR filter to,

where freq\_offset is the frequency offset from center (200Khz in this case), fine is the ID of a wx gui slider for regular tuning, and finer is the same for fine tuning. In order for the frequency display to show the proper value it was correspondingly set to a variable ID cur\_freq.



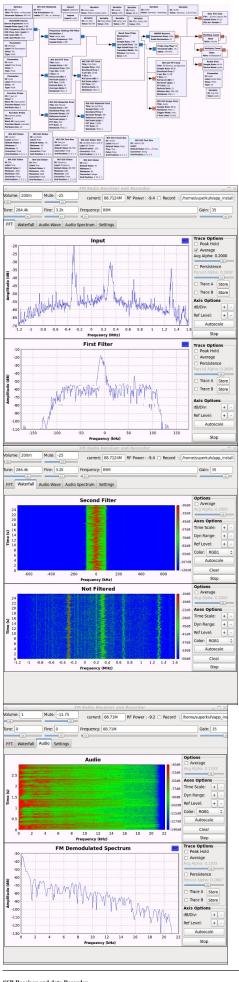
I also made the current frequency display a editable text field so you can tune, copy, and paste. There are good examples of how notebook positioning works and includes simple scripting examples for the file field. This flowchart is simple enough to learn from but includes many elements pulled out of the very complex simple for the file field. This flowchart is simple enough to learn from but includes many elements pulled out of the very complex simple fin rev from patchvonbraun. Without his explanation of the offset process I wouldn't have figured it out. All blocks are layed out by type and GUI elements in the same order as they appear when run. This should help you figure out Grid and Notebook positioning.

The sound is only "okay". I think the signal is being clipped off at the edges a little bit. I am not sure if it is required to install patchvonbraun's simple\_fm\_rev to use this, I do use some of his custom filter stuff.

Use the Waterfall for scanning through channels. Once located, look at the offset from 0 on the bottom display. Use that to set the tuning (and fine) slider and wiggle it till you get the signal crossing the band in the "Second Filter" top display. Switch to FFT view and look at the bottom "First Filter" display, use tuning and fine tuning to center the peak on the "First Filter" display. Or the other way around. It's personal preference. Ignore the noise you see at higher frequencies (900Mhz) at +0.2Mhz baseband. Although sometimes it gets folded in depending on tuning.

• superkuh's FM w/offset tuning, fine tuning, and recording (.grc)

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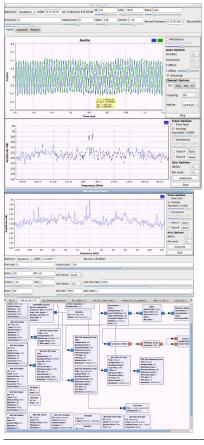
SSB Receiver and data Recorder

Created by Alexandru Csete OZ9AEC the notes say, "Simple SSB receiver prototype". This comes from the GNU Radio GRC examples repository over at https://github.com/csete/gnuradio-grc-examples/tree/master/receiver qit clone https://qithub.com/csete/qnuradio-qrc-examples.qit

Стр. 20 из 21 20.11.2020, 23:11 "/dev/null" if record == False else capture file

in the File Sink 'file' field. I also changed the GUI so it was easier to find signals. Use the saved .bin files with ssb\_rx\_play to hear. Jumping around in frequency is a lot smoother when reading from disk instead of the dongle.

- Modified OZ9AEC SSB Receiver (.grc)
   ssb\_rx\_play (.grc)



```
How to me ribdr and hackef on a 'fresh' odroid-u3 with ubuntu

# rtladr on odroid-u3
and /usr/local/bin/root-utility.sh # resize partition first
and/o apt-get update
and apt-get update
and apt-get red interpretation of the service 
  mkdir build col build (col build is for odroid-u3 weirdness. The _FILE_OFFSET_BITS is for writing >ZGB files without pipes. cnake ../ -DCMAKE_INSTALL_PREFIX=/usr -DINSTALL_UDEV_RULES=ON -DLIB_DIR=/usr/lib/arm-linux-gnueabihf -D_FILE_OFFSET_BITS=64 make sudo make install sudo su - lideonfile
Hackif one on odroid-u3, do all the non-rtladr stuff above then,
git clone git://github.com/mossmann/hackrf.git
cd hackrf/host
mkdir build
cd build
# The lib dir is for odroid-u3 weirdness. The _FILE_OFFSET_BITS is for writing >2GB files without pipes.
cmake .../ -DINSTALL_UDEV_RULES-ON -DLIB_DIR=/usr/lib/arm-linux-gnueabihf -D_FILE_OFFSET_BITS-64
make
make
make
make install
sudo au -
ldconfig
```

2020-10-26: This change is just to trigger change detection bots. Bork bork bork.

[comment on this post] Append "/@say/your message here" to the URL in the location bar and hit enter.

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